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PRACTICAL APPLICATIONS

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OF

SPECTRUM ANALYSIS

by

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PREFACE

The purpose of this book is to provide a trustworthy guide, as complete and self-contained as is possible in a single volume, for the practical spectroscopist engaged in analytical work connected with any field of study whatever. The necessary complementary volumes are tables of wave-lengths—The Massachusetts Institute of Technology Wave-Length Tables (John Wiley & Sons, Inc., New York; Chapman & Hall, Ltd., London) certainly, and if possible the Kayser/Ritschl Tabelle der Hauptlinien der Linienspektra aller Elemente nach Wellenlänge geordnet (Julius Springer, Berlin; Adam Hilger, London; 1939) as well. Pearse and Gaydon's The Identification of Molecular Spectra (Chapman & Hall) is also very desirable. Those engaged in industrial work in which saving of time is important should also keep in constant touch with the catalogues of instrument makers in order to acquaint themselves with the latest devices for particular parts of the process. No attempt has been made here to describe more than the essentials of the various forms of apparatus for producing luminosity, forming an image of the source, etc., embodied in simple but adequate equipment for all purposes except that of producing the greatest number of results in the shortest time. New devices are constantly being produced, and it would have been foolish to occupy space with descriptions of those current at the moment, which would probably have been superseded before publication.

The book is essentially a record of personal experience rather than a general conspectus of all the methods that have been used. It is an appropriate time for such a record, since I am now leaving this work after having been engaged in it for nearly thirty years, during the first half of which I was in the closest association with the late Professor A. Fowler, F.R.S., to whom I owe the foundations of my knowledge of the subject. An almost unlimited variety of problems have been presented to us at the Imperial College, nearly all of which have been parts of some wider biological, agricultural, geological, metallurgical or other investigation, and the solutions have not been separately published. It would be a pity, therefore, not to make available for others, as far as may be, the general experience gained.

The small amount of space given to quantitative spectrum analysis properly reflects, I believe, the present state of achievement in this field, but not the promise for the future. I have not thought it desirable to give more than the general principle of the many *ad hoc* methods used in particular industries—lack of space alone would have prohibited this

—but I have given what I believe must be the foundation of any practicable general method for such work. The principles of the production of spectra are now sufficiently well understood to make a theoretical approach possible, and it is most desirable that the problem should now be dealt with in this way rather than by narrowly empirical methods which, however useful they may be for their own special purposes, have no trustworthy wider application. I do not know of any previous attempt at a theoretical approach to the general problem.

While the book is in no sense a theoretical account of spectroscopy, I have tried to give sufficient theory, both of the mechanism of radiation and of the optical principles exemplified in the spectrograph, to enable the worker to form a mental picture of what is happening and to readjust his instrument for any desired end. The book should therefore be useful not only for the fortunate worker who can obtain fresh apparatus for each special purpose, but also for him who has to use the same instrument for all purposes and even to construct that instrument from parts separately acquired.

I am glad to record my thanks to various students and colleagues at the Imperial College for help with the illustrations. They are too numerous to mention personally, and of the source of some of the earlier plates I have no record. Except where special acknowledgement has been made, the illustrations were all produced at the Imperial College or have been associated with it so long that they have come to be regarded as such. For all inadvertent omissions of acknowledgement I express my regret. My thanks are also gratefully accorded to Dr. G. R. Harrison for permission to include the M.I.T. list of raies ultimes in Table V; to the Director of the Lick Observatory and Dr. K. Burns for permission to include Table II; to the Physical Society and Drs. A. Hunter and R. W. B. Pearse for permission to include the Table on p. 53; to the Institute of Physics and Dr. R. W. B. Pearse for permission to include Fig. 9; and to Messrs. Hilger & Watts, Ltd., for permission to include Plate XVIIIA, their list of raies ultimes in Table V, and the portions of the iron arc spectrum in Plates X and XI.

H. D.

London, June 1949

LIST OF PLATES

DESCRIPTION

- Note-The plates are arranged in sequence at the end of the book.
 - I-V Spectrum of raies ultimes mixture with iron arc comparison, region 6720-3700 A. (See pp. 92-5.) Two spectra of the mixture are shown, with different exposures, one on each side of the iron spectrum.
 - VI Iron arc spectrum, region 3800-3460 A.
- VII-VIII Spectrum of raies ultimes mixture with iron arc comparison, region 3560-2520 A. (See pp. 92-5.)
 - IX Iron are spectra from centre of arc and positive and negative poles, region 2845-2325 A.
 - X-XI Iron arc spectrum, region 2465-2084 A.
 - XII Absorption spectrum of nitroso-dimethyl-aniline.
 - XIII Flame, arc and spark spectra of thallium.
 - XIV (A) Arc spectrum of graphite.
 - (B) A few sequences in the CN spectrum.
 - XV Spectra showing variations in the relative intensities of VI and VII lines in the graphite arc arising from variations in the general composition of the arc gases. (See p. 88.)
 - XVI (A) Spectrum of coal-gas flame.
 - (B) Absorption spectra of neodymium chloride solutions of different strengths.
 - XVII Spectra showing danger of confusing Bi line at 3067.7 with a line in the band spectrum of OH. (See p. 99.)
 - XVIII (A) Example of standard spectra for special problem in quantitative spectrum analysis. (See p. 123.)
 - (B) Spectra showing danger of attributing lines of VII and Tim in impure graphite to BeII. (See p. 99.)
 - (C) Ghosts in grating spectrum.
 - XIX Spectra showing effect of salt in bringing out lines of impurities in graphite electrodes. (See pp. 79-80.)

PLATES I-XIX

DESCRIPTION

- I-V Spectrum of raies ultimes mixture with iron are comparison, region 6720-3700 A. (See pp. 92-5.) Two spectra of the mixture are shown, with different exposures, one on each side of the iron spectrum.
- VI Iron are spectrum, region 3800-3460 A.
- VII-VIII Spectrum of raies ultimes mixture with iron are comparison, region 3560-2520 A. (See pp. 92-5.)
 - IX Iron are spectra from centre of arc and positive and negative poles, region 2845–2325 Λ.
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 - XVIII (A) Example of standard spectra for special problem in quantitative spectrum analysis. (See p. 123.)
 - (B) Spectra showing danger of attributing lines of VII and Till in impure graphite to Bell. (See p. 99.)
 - (C) Ghosts in grating spectrum.
 - XIX Spectra showing effect of salt in bringing out lines of impurities in graphite electrodes. (See pp. 79-80.)

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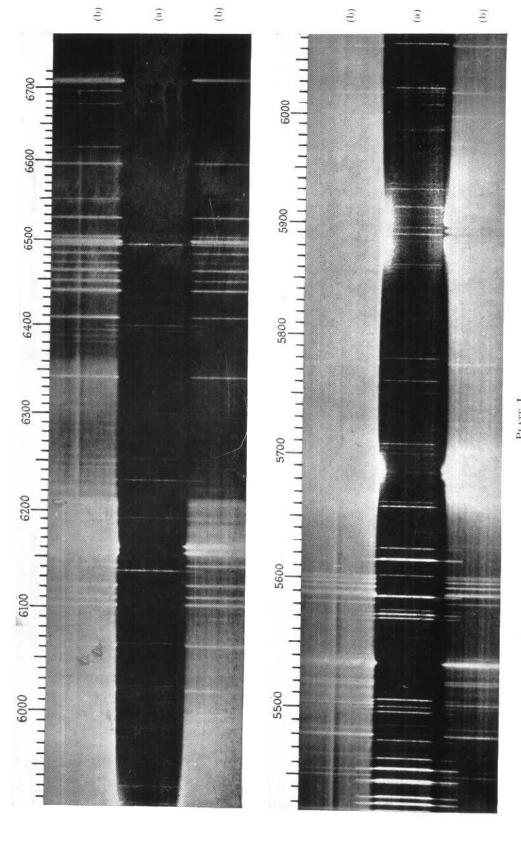
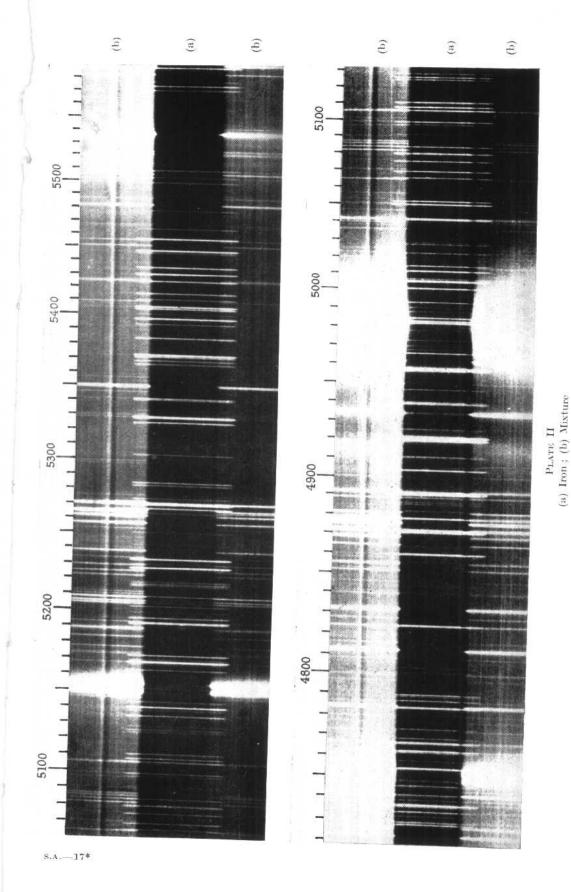
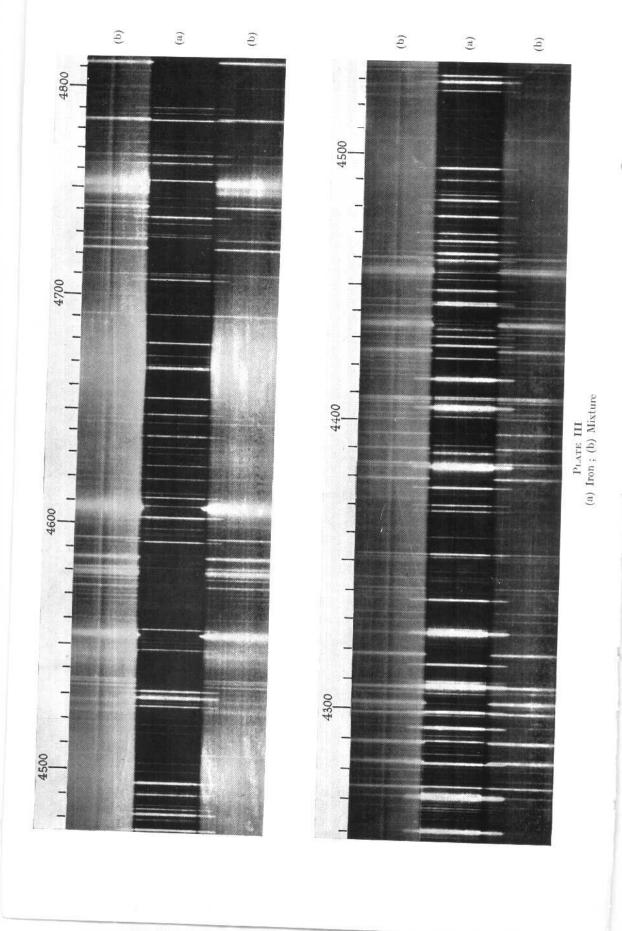


PLATE I
(a) Iron; (b) Mixture





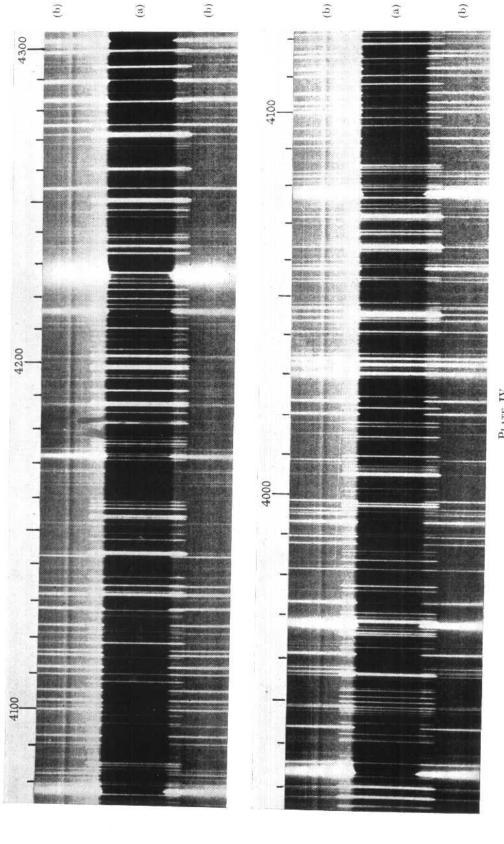
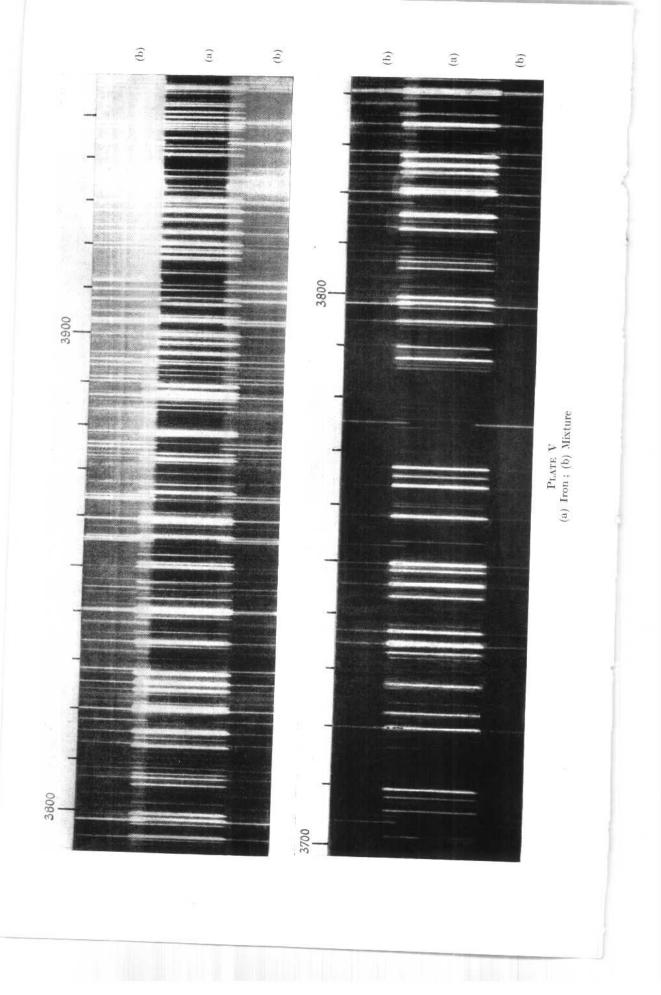
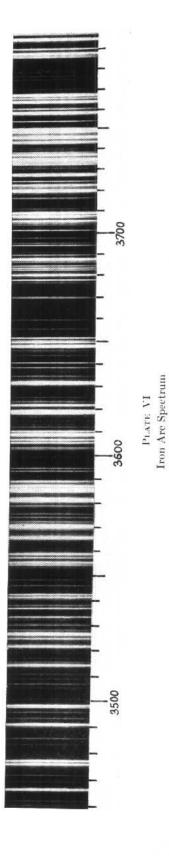


Plate IV
(a) Iron; (b) Mixture



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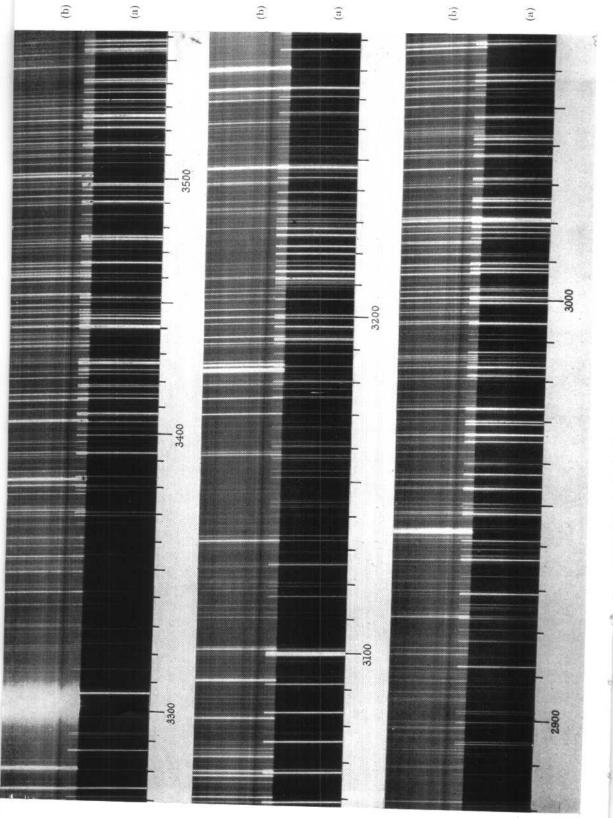


PLATE VII
(a) Iron; (b) Mixture

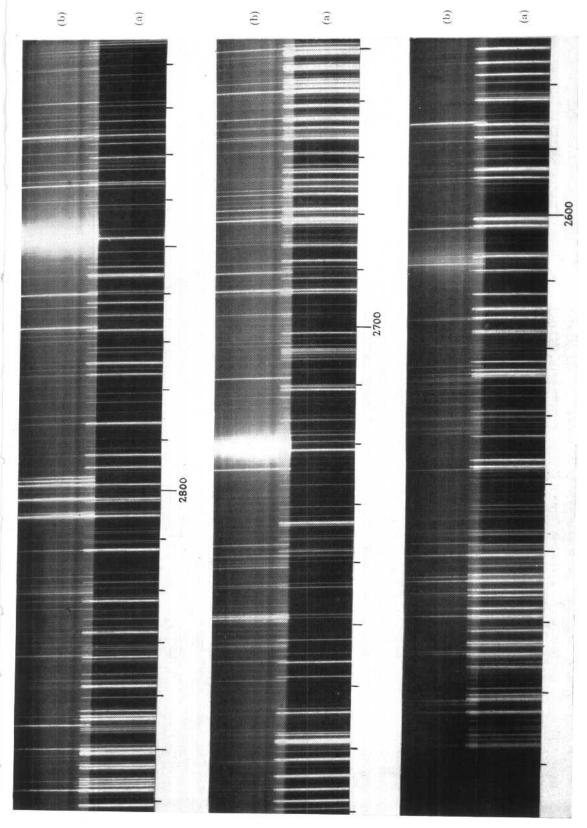
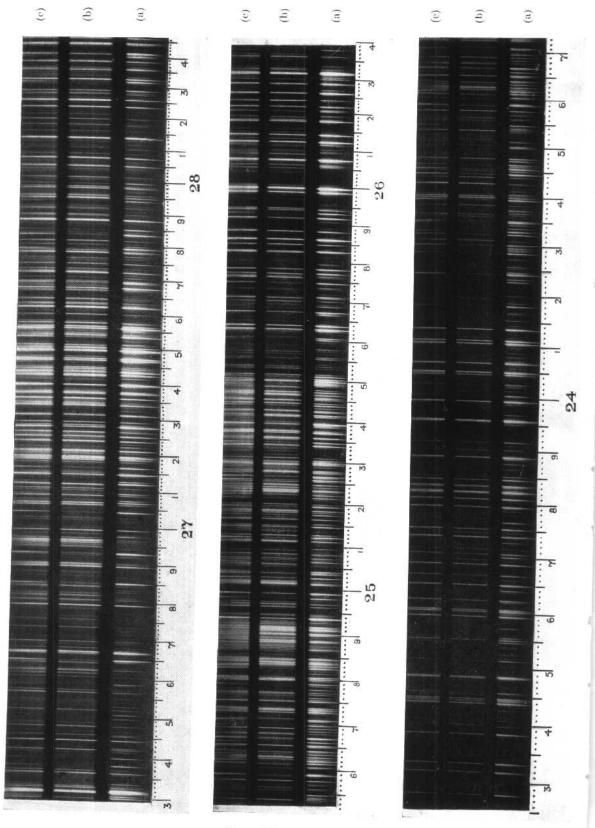


PLATE VIII
(a) Iron; (b) Mixture



 $\begin{array}{c} {\rm PLATE} \ \ IX{\rm \longrightarrow IRON} \ \ ARC \\ {\rm (a)} \ \ Positive \ \ Pole \ ; \ (b) \ \ Centre \ of \ Arc \ ; \ (c) \ \ Negative \ Pole \end{array}$

	2245.646
	2249.173
2248.855	2250.933 & .784 2251.865
2253.122	2255.859
2259.279	2260.078
2259,511	2260.860 2265.047
2264.390	2266.903 2263.047 2267.573 & .466
2267.080 2269.093	2270.858 2272.065
2271.778 2272.816	2274.087 2275.189
2277.094 2275.593	2276.023 2280.222 2277.663
2279.927 2283.299	2007 652 2283.019
2284 .083 2287 .628	2289 .032 2287 .248 2289 .032 2290 .546
2290.064	2291,117
2290.771 2292.523 2293.847 2294.926	2294 .406 2297 .787 -2298 .657
2250.520	$\begin{array}{r} 2297.107 - 2298.657 \\ 2300.139 - 2301.171 \end{array}$
2301 681	2303.422 2306.378
2303.577 2304.727 2306.164	2308.996
2306.164	2313.102
	0301 607
2320.355	2321.687
J2327.392	
2329.637 J2331.305	2332.795J
J2338.002	2343.957 2345.335
2343.492	2348.116
2344.279	2348.301
	2354.888
J2359.102	2359.997 2362.019
2360.292	2364 .825J
2366.590	2370.495 2368.594
2371 .427 2369 .454	2373.732 & .618
2374.516	2375.191 2379.273J
J2380 .757	2382.036
2383.056 2383.244	2384 . 385
J2388.625 2391.474	2389.970
2395.417	2395.623
	2399.238J
2404.429	2404.884 2406.657J
J2410.517	2411.066J
02410.021	2413.309J
2419.058	
2423.094	
	_ 2438.179
2439.743	2440.106
2439.743 2442.567	2443.870 2447.708J
	2453.472
	2457.595J
2462.178	2462.645
2402.170	2465.148J

