



PHYSICS AND CHEMISTRY OF THE SOLAR SYSTEM

second
edition

John S. Lewis



Physics and Chemistry of the Solar System

SECOND EDITION

John S. Lewis

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Physics and Chemistry of the Solar System

SECOND EDITION

This is volume 87 in the
INTERNATIONAL GEOPHYSICS SERIES
A series of monographs and textbooks
Edited by RENATA DMOWSKA, JAMES R. HOLTON, and H. THOMAS ROSSBY
A complete list of books in this series appears at the end of this volume.

Dedication

This book is dedicated to the founders of Planetary Science:
Rupert Wildt, Gerard P. Kuiper, and Harold C. Urey,
whose thoughts roamed the Solar System before spacecraft did.

Foreword

At its original conception, this book was based on the structure, scope, and philosophy of a sophomore/junior level course taught at M.I.T. by the author and Prof. Irwin I. Shapiro from 1969 to 1982. Although the content of that course varied greatly over the years in response to the vast new knowledge of the Solar System provided by modern Earth-based and spacecraft-based experimental techniques, the philosophy and level of presentation remained very much the same. The material was brought up to date in 1994 for publication in 1995, and again updated with many corrections and additions for a revised edition in 1997. This second edition was prepared in 2002 to take advantage of the many recent advances in the study of Mars and small Solar System bodies, the discovery and study of more than 100 extra-solar planets, and more mature analysis of the Galileo Orbiter and probe data on Jupiter and its large satellites.

The timing of the various editions of this book has been influenced by the erratic history of planetary exploration. During the 12 years of 1964–1973 there were 87 launches of lunar and planetary spacecraft, of which 54 were involved in the race to the Moon. In the 29 years since the end of 1973, up to the date of this edition in 2002, there have been only 36 additional launches. Both the United States and the Soviet Union experienced prolonged gaps in their lunar and planetary exploration programs: the American gap in lunar exploration extended from Explorer 49 in 1973 to the launch of

Clementine in 1994, and the Russian hiatus in lunar missions has stretched from Luna 24 in 1976 to the present. American exploration of Mars was suspended from the time of the Viking missions in 1975 until the launch of Mars Observer in 1992, and Soviet exploration of Mars, suspended after Mars 7 in 1975, did not resume until the launch of the two ill-fated Phobos spacecraft in 1988. Soviet missions to Venus ceased in 1984.

From 1982 to 1986 there was a gap in the acquisition of planetary data by American spacecraft. This drought was interrupted in 1986 by the Voyager 2 Uranus flyby and by five spacecraft encounters with Halley's comet (two Soviet, two Japanese, and one from the European Space Agency), but the drought again resumed until it was broken by the Voyager 2 Neptune encounter and the Soviet Phobos missions in 1989 and the Magellan mission to Venus in 1990. The launch of the Galileo Orbiter and probe to Jupiter, long scheduled for 1986, was severely delayed by the explosion of the space shuttle orbiter Challenger, the resulting 2-year grounding of the entire shuttle fleet, and the subsequent cancellation of the high-energy Centaur G' upper stage intended for launching heavy planetary missions from the shuttle. The European-American Ulysses solar mission, which was not instrumented for intensive planetary studies, flew by Jupiter in February 1992, returning only data on its magnetic and charged-particle environment. The arrival of Galileo at Jupiter, the Galileo Probe entry into

Jupiter's atmosphere in December 1995, the lengthy Galileo Orbiter survey of the Jovian system, and the resumption of small Mars missions (Pathfinder, Mars Global Surveyor, etc.) by the United States have combined with a flood of space-based (Galileo, Near-Earth Asteroid Rendezvous) and Earth-based observations of near-Earth asteroids and Belt asteroids, and intensive Earth-based study of comets, Centaurs, small icy satellites, and trans-Neptunian objects and the highly successful search for dark companions of nearby stars to reinvigorate the planetary sciences. This new resurgence of planetary exploration, with little prospect of Russian participation, has been helped by the active involvement of Japan's NASDA and the European Space Agency in planning and flying unmanned missions to the Moon, Mars, and Venus. The infusion of new data resulting from these several programs creates the necessity of revising this book.

In this book, as in that Planetary Physics and Chemistry course in which it was first conceived, I shall assume that the reader has completed 1 year of university-level mathematics, chemistry, and physics. The book is aimed at several distinct audiences: first, the upper-division science major who wants an up-to-date appreciation of the present state of the planetary sciences for "cultural" purposes; second, the first-year graduate student from any of several undergraduate disciplines who intends to take graduate courses in specialized areas of planetary sciences; and third, the practicing Ph.D. scientist with training in physics, chemistry, geology, astronomy, meteorology, biology, etc., who has a highly specialized knowledge of some portion of this material, but has not had the opportunity to study the broad context within which that specialty might be applied to current problems in this field.

This volume does not closely approximate the level and scope of any previous book. The most familiar texts on the planetary sciences are *Exploration of the Solar System*, by William J. Kaufmann, III (Macmillan, New York, 1978 and later), a nonmathematical survey of the history of planetary exploration; *Moons and Planets*, by William K. Hartmann (Wadsworth, Belmont, California, 1972; 1983; 1993), a scientific tour of the Solar System with high-school-level mathematical content; and *Meteorites and the Origin of Planets*, by John A. Wood (McGraw-Hill, New York, 1968), a fine qualitative introduction that is similarly sparing of mathematics and physics. Several other nonmathematical texts are available, including *Introduction to the Solar System*, by Jeffrey K. Wagner (Saunders, Philadelphia, 1991), *Exploring the Planets*, by W. Kenneth Hamblin and Eric H. Christiansen (Macmillan, New York, 1990), *The Space-Age Solar System*, by Joseph F. Baugher (J. Wiley, New York, 1988), and *The Planetary System*, by

planetary scientists David Morrison and Tobias Owen (Addison-Wesley, Reading, Massachusetts, 1988).

Another book, comparable in mathematical level to the present text, is *Worlds Apart*, by Guy J. Consolmagno, S. J., and Martha W. Schaefer (Prentice Hall, Englewood Cliffs, New Jersey, 1994). Though much less detailed than the present work, it is well written and appropriate for a one-semester introductory course on planetary science for science majors. The scope of the present text is broader, and the level higher, than any of these books.

As presently structured, this book is a broad survey of the Solar System suitable for reference use or as background reading for any course in Solar System science. The text may for convenience be divided into three parts. The first of these parts contains Chapter I (Introduction), Chapter II (Astronomical Perspective), Chapter III (General Description of the Solar System), and Chapter IV (The Sun and the Solar Nebula). This first part could be called "General Properties and Environment of our Planetary System." It is roughly equivalent to a brief introductory astronomy book emphasizing the concerns of planetary scientists rather than stellar or galactic astronomers. The second part contains Chapter V (The Major Planets), Chapter VI (Pluto and the Icy Satellites of the Outer Planets), Chapter VII (Comets and Meteors), and Chapter VIII (Meteorites and Asteroids), and might fairly be entitled "The Solar System beyond Mars." The third and final part comprises Chapter IX (The Airless Rocky Bodies: Io, Phobos, Deimos, the Moon, and Mercury), Chapter X (The Terrestrial Planets: Mars, Venus, and Earth), Chapter XI (Planets and Life around Other Stars), and Chapter XII (Future Prospects). This part could be called "The Inner Solar System."

Using this volume as a textbook, a planetary sciences course taught in a trimester setting could use one part each term. In a two-semester program, either an inner solar system emphasis course (parts 1 and 3) or an outer solar system course (parts 1 and 2) could be taught. The most ambitious and intensive program, and the most similar to the way the course was structured at M.I.T., would be to teach parts 2 and 3 in two semesters, reserving most of the material in part 1 for use as reference reading rather than as lecture material.

This book is written in appreciation of the approximately 350 students who took the course at M.I.T., and who unanimously and vocally deplored the lack of a textbook for it. These students included both Consolmagno and Schaefer as cited above. I extend my particular thanks to Irwin Shapiro for his many years of cheerful, devoted, always stimulating, and sometimes hilarious collaboration on our course,

and for his generous offer to allow me to write “his” half of the text as well as “mine.” I am also pleased to acknowledge the helpful comments and suggestions of dozens of my colleagues, but with special thanks

reserved for Jeremy Tatum of the University of Victoria, whose detailed comments and physicist’s perspective have been invaluable in the preparation of this second edition.

Contents

Foreword xi

I Introduction

Nature and Scope of Planetary Science 1
Guide to the Literature 3
Numbers in Science 4
Dimensions and Units 5
Exercises 6

II Astronomical Perspective

Introduction 7
Distance Scales in the Universe 7
The Big Bang 10
Limitations on Big Bang Nucleosynthesis 14
Galaxy and Star Formation 15
Structure and Classification of Galaxies 16
Classification of Stars 18
Stellar Evolution 25
Star Clusters 27
Stellar Origins 29

Outline of Star Formation 33
Stellar Explosions and Nucleosynthesis 34
Nuclear Cosmochronology 43
Exercises 47

III General Description of the Solar System

Introduction 50
The Sun 50
Orbits of the Planets 52
Changes in Orbital Motion 57
Properties of the Planets 58
Mass and Angular Momentum Distribution 59
Satellites 63
Asteroids 69
Comets 71
Meteors 72
Meteorites 72
Cosmic Dust 73
Cosmic Rays 73
Planetary Science in the Space Age 74

Summary	76
Exercises	76

IV The Sun and the Solar Nebula

Introduction	77
Energy Production in the Sun	77
Energy Transport in the Sun	79
Internal Structure of the Sun	83
Surface of the Sun	84
The Chromosphere	87
The Corona	88
Discovery of the Solar Wind	90
Radio Wave Propagation in Space Plasmas	91
The Solar Wind	92
Chemistry of Solar Material	96
Ionization	97
Dissociation and Molecule Formation	100
Hydrogen and the Rare Gases	101
Oxygen, Carbon, and Nitrogen	102
Magnesium and Silicon	105
Iron	106
Sulfur	107
Aluminum and Calcium	108
Sodium and Potassium	109
Nickel and Cobalt	110
Phosphorus and the Halogens	111
Geochemical Classification of the Elements	111
The Chemistry of Rapid Accretion	116
Kinetic Inhibition	117
Mass and Density of the Solar Nebula	118
Thermal Opacity in the Solar Nebula	121
Dust Opacity	129
Thermal Structure of the Nebula	131
Turbulence and Dust Sedimentation	134
Accretion of Rocks, Planetesimals, and Planets	136
Gas Capture from the Solar Nebula	138
The T Tauri Phase	141
Thermal History of the Early Solar System	143
Exercises	144

V The Major Planets

Introduction	147
Interiors of Jupiter and Saturn: Data	148
Isothermal Interior Models of Jupiter and Saturn	151
Thermal Models of Jupiter and Saturn	154
The Atmospheres of Jupiter and Saturn: Observed Composition	156

Tropospheric Composition and Structure: Theory	159
Cloud Condensation in the $\text{NH}_3\text{--H}_2\text{O--H}_2\text{S}$ System	165
Cloud Physics on the Jovian Planets	174
Galileo Perspectives on Jovian Clouds	179
Ion Production in the Jovian Atmosphere	180
Visible and Infrared Radiative Transfer	183
Horizontal Structure and Atmospheric Circulation	187
Photochemistry and Aeronomy	200
The Jovian Thermosphere	217
Radiophysics and Magnetospheres of Jupiter and Saturn	218
The Interiors of Uranus and Neptune	229
Atmospheres of Uranus and Neptune	238
Perspectives	247
Exercises	247

VI Pluto and the Icy Satellites of the Outer Planets

Introduction	252
Surfaces of Icy Satellites	253
Eclipse Radiometry	256
Surface Temperatures	257
Surface Morphology of the Galilean Satellites	258
Density and Composition of Icy Satellites	265
Internal Thermal Structure of Galilean Satellites	267
Dynamical Interactions of the Galilean Satellites	272
Thermal and Tectonic Evolution of Icy Satellites	275
Minor Satellites of Jupiter	278
Planetary Rings	280
Titan	289
The Intermediate-Sized Saturnian Satellites	293
Minor Satellites of Saturn	296
Satellites of Uranus	299
Satellites of Neptune	303
The Pluto–Charon System	308
The Neptune–Pluto Resonance	311
Spacecraft Exploration	311
Exercises	312

VII Comets and Meteors

Historical Perspectives	317
Nature and Nomenclature of Comets	319

Cometary Orbits	321
Heating by Passing Stars	325
Evaporation and Nongravitational Forces	326
The Nucleus and Coma of P/Halley	328
Chemistry and Photochemistry of Water	328
Further Chemical Processes in the Coma and Tail	332
Behavior of Small Particles	333
Dynamical Behavior of Dust in Space	334
Meteors	336
Cometary Fireballs	343
Cometary Impacts on Jupiter	344
Exercises	347

VIII Meteorites and Asteroids

Introduction	350
Introduction to Meteorites	350
Meteorite Orbits	353
Phenomena of Fall	355
Physical Properties of Meteorites	358
Meteorite Minerals	362
Taxonomy and Composition of Chondrites	362
Metamorphic Grades of Chondrites	367
Taxonomy and Composition of Achondrites	369
Taxonomy and Composition of Stony-Irons	371
Taxonomy and Composition of Irons	372
Isotopic Composition of Meteorites	375
Genetic Relationships between Meteorite Classes	382
Introduction to Asteroids	384
Asteroid Orbits	386
Stability of Trojan and Plutino Orbits	389
Sizes, Shapes, and Albedos of Asteroids	391
Masses and Densities of Asteroids	393
Photometry and Spectroscopy of Asteroids	394
Thermal Evolution of Asteroids	401
Dynamical Evolution of the Asteroid Belt	406
Centaurs and Trans-Neptunian Objects	409
Relationships among Asteroids, Meteorites, and Comets	412
Radar Observations of Near-Earth Asteroids	415
Asteroid Resources	416
Exercises	419

IX The Airless Rocky Bodies: Io, Phobos, Deimos, the Moon, and Mercury

Introduction	424
Orbits and Physical Structure of Phobos and Deimos	426

Io: General Properties	430
Io: Surface Processes	430
Io: Internal Energy Sources	432
Io: Geology	433
Io: Atmospheric and Volcanic Gases	435
Io: Escape and the Plasma Torus	437
Io: Genetic Relationships	438
Impact Cratering	438
Motions of the Moon	443
Physical Properties of the Moon	445
Elemental Composition of the Moon's Surface	445
Lunar Rock Types	447
Lunar Minerals	449
Lunar Elemental Abundance Patterns	451
Geology of the Moon	451
Geophysics of the Moon	452
History of the Earth-Moon System	456
Origin and Internal Evolution of the Moon	458
Solar Wind Interaction with the Moon and Mercury	460
The Planet Mercury	461
Motions of Mercury	461
Composition and Structure of Mercury	462
Noncrater Geology of Mercury	463
Geophysics of Mercury	463
Atmospheres of Mercury and the Moon	468
Polar Deposits on Mercury and the Moon	469
Unfinished Business	472
Exercises	474

X The Terrestrial Planets: Mars, Venus, and Earth

Introduction	477
Mars	478
Motions of Mars	479
Density and Figure of Mars	479
Geophysical Data on Mars	481
Gravity and Tectonics of Mars	483
Geology of Mars	483
Surface Composition	496
Viking Lander Investigations	503
The Shergottite, Nakhilite, and Chassignite Meteorites	505
Atmospheric Structure	508
Atmospheric Circulation	509
Atmospheric Composition	510
Photochemical Stability and Atmospheric Escape	513
Explosive Blowoff	519
Origin and Evolution of the Atmosphere	519

Organic Matter and the Origin of Life	522
Venus	524
Motions and Dynamics of Venus	526
Geophysical Data on Venus	526
Geology of Venus	528
Venus: Atmospheric Structure and Motions	534
Venus: Atmospheric Composition	537
Venus: Atmosphere–Lithosphere Interactions	539
Venus: Photochemistry and Aeronomy	543
Venus: Atmospheric Escape	547
Venus: Planetary Evolution	549
Earth	550
Earth: Motions	551
Earth: Internal Structure	552
Earth: Magnetic Field and Magnetosphere	554
Earth: Surface Geology	554
Earth: Early Geological History	557
Earth: Biological History	559
Earth: Geochemistry and Petrology	563
Weathering in the Rock Cycle	566
Earth: Atmospheric Composition and Cycles	568
Radiocarbon Dating	573
Stable Isotope Climate Records	574
Photochemistry and Aeronomy	575
Escape and Infall	575
Climate History, Polar Ice, and Ice Ages	579
Life: Origins	582
Life: Stability of the Biosphere	587
Exercises	588

XI Planets and Life around Other Stars

Chemical and Physical Prerequisites of Life	592
The Planetary Environment	595
The Stellar Environment	597
Brown Dwarfs	600
The Search for Planets of Other Stars	603
The Search for Extraterrestrial Intelligence	606
Exercises	608

XII Future Prospects

Mercury	611
Venus	612
Earth's Moon	612

Mars	613
Asteroids	614
Jupiter	615
Saturn, Uranus, and Neptune	615
Pluto	615
Comets	616
Beyond the Solar System	616

Appendix I: Equilibrium Thermodynamics 621

Heat and Work	621
Adiabatic Processes and Entropy	622
Useful Work and the Gibbs Free Energy	623
Chemical Equilibrium	623
Exact and Complete Differentials	624
The Maxwell Relations	625

Appendix II: Absorption and Emission of Radiation by Quantum Oscillators 626

Appendix III: Exploration of the Solar System 629

Appendix IV: Basic Physical Constants 634

Appendix V: Gravity Fields 635

Suggested Readings

Introduction	637
Chapter I–Introduction	637
Chapter II–Astronomical Perspective	637
Chapter III–General Description of the Solar System	638
Chapter IV–The Sun and the Solar Nebula	638
Chapter V–The Major Planets	638
Chapter VI–Pluto and the Icy Satellites of the Outer Planets	639
Chapter VII–Comets and Meteors	639
Chapter VIII–Meteorites and Asteroids	639
Chapter IX–The Airless Rocky Bodies: Io, Phobos, Deimos, the Moon, and Mercury	640
Chapter X–The Terrestrial Planets: Mars, Venus, and Earth	640
Chapter XI–Planets and Life around Other Stars	641
Chapter XII–Future Prospects	642

Index 643

I. Introduction

Nature and Scope of the Planetary Sciences

When asked in an interview to give his viewpoint on the frontiers of science, the famous physicist Victor Weisskopf commented that the most exciting prospects fell into two categories, the frontier of size and the frontier of complexity. A host of examples come to mind: cosmology, particle physics, and quantum field theory are clearly examples of the extremes of scale, and clearly among the most exciting frontiers of science. Biology, ecology, and planetary sciences are equally good examples of the frontier of complexity.

When we peruse the essential literature of planetary science, we find that we must, over and over again, come face to face with these same extremes. First, we are concerned with the origin and nuclear and chemical evolution of matter, from its earliest manifestation as elementary particles through the appearance of nuclei, atoms, molecules, minerals, and organic matter. Second, on the cosmic scale, the origin, evolution, and fate of the Universe emerge as themes. Third, we are confronted with the problem of understanding the origin and development of life. In each case, we are brought face to face with the spontaneous rise of extreme complexity out of extreme simplicity, and with the intimate inter-relationship of the infinitesimally small and the ultimately large.

Further, our past attempts at addressing these three great problems have shown us that they are remarkably intertwined. The very issue of the origin of life is inextricably tied up with the chemistry of interstellar clouds, the life cycles of stars, the formation of planets, the thermal and outgassing history of planetary bodies, and the involvement of geochemical processes in the origin of organic matter. The connection between life and planetary environments is so fundamental that it has been given institutional recognition: it is not widely known outside the field, but research on the origin of life in the United States is a mandate of the National Aeronautics and Space Administration.

Wherever we begin our scientific pilgrimage throughout the vast range of modern science, we find ourselves forced to adopt ever broader definitions of our field of interest. We must incorporate problems not only on the frontier of complexity, but also from both extreme frontiers of scale. In this way, we are compelled to trespass across many hallowed disciplinary boundaries.

Further, as we seek an evolutionary account of the emergence of complexity from simplicity, we become able to see more clearly the threads that lead from one science to another. It is as if the phenomena of extreme scale in physics existed for the express purpose of providing a rationale for the existence of astronomy.

The other disciplines evolve logically from cosmic events:

The astronomical Universe, through the agency of nuclear reactions inside stars and supernova explosions, populates space with atoms of heavy elements, which are the basis of chemistry.

The course of spontaneous chemical evolution of interstellar matter produces both mineral grains and organic molecules, giving rise to geochemistry and organic chemistry.

Solid particles accrete to form large planetary bodies, and give us geology.

Radioactive elements formed in stellar explosions are incorporated into these planets, giving life to geophysics.

Melting, density-dependent differentiation, and outgassing take place, and atmospheres and oceans appear: petrology, meteorology, and oceanography become possible.

Organic matter is formed, accumulated, concentrated, and processed on planetary surfaces, and biology is born.

Planetary science may then be seen as the bridge between the very simple early Universe and the full complexity of the present Earth. Although it partakes of the excitement of all of these many fields, it belongs to none of them. It is the best example of what an interdisciplinary science should be: it serves as a unifying influence by helping to dissolve artificial disciplinary boundaries, and gives a depth and vibrancy to the treatment of evolutionary issues in nature that transcends the concerns and the competence of any one of the parent sciences. But there is more: planetary science is centrally concerned with the evolutionary process, and hence with people's intuitive notion of "how things work." There is as much here to unlearn as there is to learn.

We, at the turn of the millennium, still live under the shadow of the clockwork, mechanistic world view formulated by Sir Isaac Newton in the 17th century. Even the education of scientists is dedicated first and foremost to the inculcation of attitudes and values that are archaic, dating as they do from Newton's era: viewpoints that must be unlearned after sophomore year. We are first led to expect that the full and precise truth about nature may be extracted by scientific measurements; that the laws of nature are fully knowable from the analysis of experimental results; that it is possible to predict the entire course of future events if, at one moment, we should have sufficiently detailed information about the distribution and motion of matter. Quantum mechanics and relativity are later taught to us as a superstructure on Newtonian physics, not vice versa. We must internally turn our education upside down to accommodate a universe that is fundamentally quantum-mechanical,

chaotic, and relativistic, within which our "normal" world is only a special case.

All of these issues come to bear on the central question of the evolution of the cosmos and its constituent parts. Most of us have had a sufficient introduction to equilibrium thermodynamics to know that systems spontaneously relax to highly random, uninteresting states with minimum potential energy and maximum entropy. These are the classical conclusions of J. Willard Gibbs in the 19th century. But very few of us are ever privileged to hear about the development of nonequilibrium thermodynamics in the 20th century, with its treatment of stable dissipative structures, least production of entropy, and systems far removed from thermodynamic equilibrium. Think of it: systems slightly perturbed from equilibrium spontaneously relax to the dullest conceivable state, whereas systems far from equilibrium spontaneously organize themselves into structures optimized for the minimization of disorder and the maximization of information content!

It is no wonder that the whole idea of evolution is so magical and counterintuitive to so many people, and that the critics of science so frequently are able to defend their positions by quoting the science of an earlier century. We often hear expressed the idea that the spontaneous rise of life is as improbable as that a printshop explosion (or an incalculable army of monkeys laboring at typewriters) might accidentally produce an encyclopedia. But have we ever heard that this argument is obsolete nonsense, discredited by the scientific progress of the 20th century? Sadly, there is a gap of a century between the scientific world view taught in our schools and the hard-won insights of researchers on the present forefront of knowledge. The great majority of all people never learn more than the rudiments of Newtonian theory, and hence are left unequipped by their education to deal with popular accounts of modern science, which at every interesting turn is strikingly non-Newtonian. News from the world of science is, quite simply, alien to them. The message of modern science, that the Universe works more like a human being than like a mechanical wind-up toy, is wholly lost to them. Yet it is precisely the fundamental issues of how things work and how we came to be, what we are and what may become of us, that are of greatest human interest. The "modern" artist or writer of the 20th century often asserted modernity by preaching the sterility of the Universe and the alienation of the individual from the world. But this supposed alienation of the individual from the Universe is, to a modern scientist, an obsolete and discredited notion.

The problems of evolutionary change and ultimate origins are not new concerns. Far from being the private domain of modern science, they have long been among the chief philosophical concerns of mankind. Astronomy

and astrology were the parents of modern science. The earliest human records attest to mankind's perpetual fascination with origins:

Who knows for certain and can clearly state
Where this creation was born, and whence it came?
The *devas* were born after this creation,
So who knows from whence it arose?

No one knows where creation comes from
Or whether it was or was not made:
Only He who views it from highest heaven knows;
Surely He knows, for who can know if He does not?

Rigveda X 129.6–7
Circa 3000 BC

Such an attitude, reflective of curiosity, inquiry, and suspended belief, is admirably modern. But today, in light of the exploration of the Solar System, we need no longer regard our origins as complete mysteries. We can now use the observational and theoretical tools of modern science to test rival theories for their faithfulness to the way the Universe really is. Some theories, when tested by the scientific method, are found to give inaccurate or even blatantly wrong descriptions of reality and must be abandoned. Other theories seem to be very reliable guides to how nature works and are retained because of their usefulness. When new data arise, theories may need to be modified or abandoned. Scientific theories are not absolute truth and are not dogma: they are our best approximation of truth at the moment. Unlike dogma, scientific theories cannot survive very long without confronting and accommodating the observed facts. The scientific theories of today are secondary to observations in that they are invented—and modified—by human beings in order to explain observed facts. They are the result of an evolutionary process, in which the “most fit” theories (those that best explain our observations) survive. In planetary science, that process has been driven in recent years in part by the discovery and study of several new classes of bodies both within our Solar System and elsewhere. It is the great strength of science (not, as some allege, its weakness) that it adapts, modifies, and overthrows its theories to accommodate these new realities. Our plan of study of the Solar System mirrors this reality.

This book will begin with what little we presently know with confidence about the earliest history of the Universe, and trace the evolution of matter and its constructs up to the time of the takeover of regulatory processes on Earth by the biosphere. We introduce the essential contributions of the various sciences in the order in which they were invoked by nature, and build complexity upon complexity stepwise. Otherwise, we might be so overawed by the complexity of Earth, our first view of nature, that we might despair of ever gaining any understanding at all.

This approach should also dispel the notion that we are about to understand everything. It is quite enough to see that there are untold vistas for exploration, and more than enough of the Real to challenge our most brilliant intellects and most penetrating intuitions.

Let us approach the subject matter covered herein with the attitude that there are a number of fundamental principles of nature, of universal scope, that allow and force the evolutionary process. With our senses at the most alert, willing to entertain the possibility of a host of hypotheses, and determined to subject all theories and observations alike to close scrutiny, we are challenged to grasp the significance of what we see. Let us cultivate the attitude that the ultimate purpose of the planetary sciences is to uncover enough of the blueprints of the processes of evolution so that we will be able to design, build, and operate our own planetary system.

Like it or not, we are assuming responsibility for the continued stability and habitability of at least one planet. The scale of human endeavor has now become so large that our wastes are, quite inadvertently, becoming major factors in global balances and cycles. Soon our scope may be the whole Solar System. The responsible exercise of our newly acquired powers demands an understanding and consciousness superior to that which we have heretofore exhibited. Now is the time for us to learn how planets work.

Guide to the Literature

It is difficult, as we have seen above, to draw a tidy line around a particular portion of the scientific literature and proclaim all that lies outside that line to be irrelevant. Still, there are certain journals that are more frequently used and cited by practitioners of planetary science. Every student should be aware both of these journals and the powerful abstracting and citation services now available.

Astronomical observations, especially positional measurements, orbit determinations, and the like that are carried out using Earth-based optical, radio, and radar techniques, are often published in the *Astronomical Journal* (AJ). Infrared spectroscopic and radiometric observations and a broad range of theoretical topics often appear in the *Astrophysical Journal* (ApJ). The most important journals devoted to planetary science in the broad sense are *Icarus* and the *Journal of Geophysical Research* (usually called JGR). Two journals are devoted to relatively quick publication of short related papers: *Geophysical Research Letters* (GRL) and *Earth and Planetary Science Letters* (EPSL). Two general-purpose wide-circulation journals also frequently publish planetary science papers, including special issues on

