# Radiation Heat Transfer Notes

D. K. Edwards

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# **RADIATION HEAT TRANSFER NOTES**

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### **PREFACE**

These Notes are intended for a 40-lecture-hour course in radiation heat transfer. The emphasis is on developing the information and skills to formulate and solve engineering heat transfer problems. The Notes are in two main parts. After an introduction to the fundamental concepts of spectral and total intensity and flux and fractional functions, the first part develops the subject of surface-to-surface radiation heat transfer. Surface radiation characteristics needed to incorporate radiation heat transfer rates into a first law of thermodynamics heat balance are defined and related to fundamental physical properties where appropriate via electromagnetic theory. The experimental techniques developed to measure the characteristics are briefly described. Then the engineering enclosure problem is addressed. Radiosity-irradiation formulations, the network method, the mirror-image concept, and the powerful Monte Carlo algorithm are described. The second part considers the subject of radiation transfer in a participating medium. Gas radiation properties are defined, and their measurement described. The engineering enclosure problem with a well-stirred and thus homogeneous absorbing-emitting medium is treated using radiosity-irradiation and network representations. The Monte Carlo algorithm is extended to cover an absorbing-emitting-scattering medium that can be homogeneous or inhomogeneous. Geometrical and spectral complexities and simplifications are considered, primarily for the slab. Differential formulations are briefly described. Narrow- and wide-band scaling for inhomogeneous molecular gases complete the second part concerned with gas radiation. The Notes conclude with a short section on heat transfer by radiation and conduction or convection.

To keep within the confines of a 40-hour course, the Notes keep to the engineering radiation heat transfer discipline. Even so, there is an attempt to develop two levels of problem-solving skills. The first level, of great importance to practicing engineers, is what may be termed finite-element thermal system analysis. This skill is emphasized first in the development, in Sections 3 and 6, and in the early portion of Section 8. The second level is differential-element or so-called exact analysis. While this level is almost totally absent from Sections 3 and 6, the bulk of Sections 7 and 8 are devoted to it.

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D. K. Edwards

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# PRINCIPAL SYMBOLS

| a                                | length between surface roughness asperities, $\mu\text{m}$   |  |  |
|----------------------------------|--|--|--|
| A,A <sub>i</sub> ,A <sub>j</sub> | area, $m^2$  |  |  |
| A <sub>c</sub>                   | cross-sectional area, m <sup>2</sup>   |  |  |
| $A_{\mathbf{k}}$                 | band absorption of kth band, cm <sup>-1</sup>  |  |  |
| $A_{\mathbf{k}}^{\star}$         | band absorptance at kth band, dimensionless  |  |  |
| A <sup>+</sup>                   | Van Driest's constant, 26  |  |  |
| b                                | thickness, m; also Mei and Squire constant, 3.4  |  |  |
| В                                | black body heat flux, W/m², or the spectral black body heat flux, W/m² cm⁻¹ or W/m² µm; also the rotational constant, cm⁻¹ |  |  |
| B                                | magnetic induction   |  |  |
| c                                | velocity of light, $2.997925x10^8$ m/s   |  |  |
| c <sub>1</sub>                   | first radiation constant, $3.7415 \times 10^{-16} \text{ J m}^2/\text{s}$  |  |  |
| c <sub>2</sub>                   | second radiation constant, 1.4388 cm K   |  |  |
| c <sub>f</sub>                   | skin friction coefficient  |  |  |
| c <sub>p</sub>                   | specific heat at constant pressure, J/kg K   |  |  |
| C*                               | channel emissivity, dimensionless  |  |  |
| d                                | spectral line spacing, cm <sup>-1</sup>  |  |  |
| D                                | diameter, m; also optical collision diameter   |  |  |
| $\bar{\mathbf{D}}$               | electric induction   |  |  |
| E                                | energy, J or eV  |  |  |
| E <sub>n</sub>                   | exponential integral of order n  |  |  |
| Ē                                | electric field strength  |  |  |
| f                                | function of fractional function; also friction factor  |  |  |
| f <sub>e</sub>                   | external fractional function, dimensionless  |  |  |

| $\mathbf{f_i}$      | internal fractional function, dimensionless  |
|---------------------|--|
| F                   | collision frequency, $s^{-1}$ ; also entry length factor   |
| F <sub>i-j</sub>    | shape factor from area i to area j, dimensionless  |
| $\mathcal{F}_{i-j}$ | transfer factor from area i to area j, dimensionless   |
| g                   | statistical weight   |
| G                   | normal irradiation of a beam, $f \operatorname{Id}\Omega$ , $W/m^2$  |
| G <sub>i-j</sub>    | area-shape-factor product, $A_iF_{i-j}$ , $m^2$  |
| g <sub>i−j</sub>    | area-transfer-factor product A $\mathscr{F}_i$ ; also the areato-volume or volume-to-volume quantity, $\mathbf{m}^2$ |
| h                   | Planck's constant, 6.6256x10 <sup>-34</sup> Js   |
| h <sub>c</sub>      | convective heat transfer coefficient   |
| $^{ m h}{}_{ m r}$  | radiative heat transfer coefficient  |
| ĥ                   | specific enthalpy, J/kg  |
| <u> </u>            | magnetic field strength  |
| Ι                   | intensity, W/m $^2$ sr; also the spectral intensity, W/m $^2$ cm $^{-1}$ sr or W/m $^2$ $\mu m$ sr                   |
| I <sub>b</sub>      | black body intensity (units as above)  |
| j                   | electric current density   |
| k                   | absorptive index; also Boltzmann's constant, $1.38054x10^{-23}$ J/K  |
| k <sub>a</sub>      | absorption coefficient, $m^{-1}$   |
| <sup>k</sup> e      | extinction coefficient, $m^{-1}$   |
| k <sub>s</sub>      | scattering coefficient, $m^{-1}$   |
| K                   | thermal conductivity, W/m K; also the second moment of the radiant intensity, also von Karman's constant, $0.4$      |
| $K_{R}$             | radiation conductivity, W/m K  |
| K*                  | radiation kernel   |

L length, m mean beam length, m  $^{\rm L}$ mb geometric mean beam length, m Lmbg mass, kg m mass flow rate, kg/s m refractive index n radiation conductivity to molecular conductivity N ratio, dimensionless number density. m<sup>-3</sup> N Nu Nusselt number pressure, N/m<sup>2</sup> or atm P Pperimeter, m Pe dimensionless equivalent line broadening pressure P; partial pressure of species i  $Pr_{m}$ Prandtl number Prt turbulent Prandtl number heat flux, W/m<sup>2</sup> q radiosity, W/m<sup>2</sup> q<sup>†</sup>

q heat flux, W/m

q<sup>†</sup> radiosity, W/m<sup>2</sup>

q<sup>-</sup> irradiation, W/m<sup>2</sup>

Q heat energy, J

Q heat flow, W

Q<sub>a</sub> absorption efficiency

Q<sub>e</sub> extinction efficiency

Q<sub>s</sub> scattering efficiency

electrical resistivity, ohm cm r  $\mathbf{r}_{\mathtt{WL}}$ wall-layer thickness-to-hydraulic diameter ratio, dimensionless R radius, m R+ turbulent Reynolds number Re Reynolds number based upon hydraulic diameter slant path length, m S S integrated line intensity; also mean radiant intensity t time, s; also optical depth, dimensionless optical depth at band head, dimensionless  $t_{H}$ T temperature, K Te environmental temperature or equivalent mean temperature, K Tg gas temperature, K mean radiant temperature, K  $T_r$ Tw wall temperature, K u sec  $\theta$ ; also cos  $\theta$ ; also velocity in x-direction. m/s U unit step function; also average velocity in xdirection, m/s cos θ ν V detector signal, volts length coordinate, m; also mole fraction Х density-path-length product, kg/m<sup>2</sup> or g/m<sup>2</sup> X length coordinate, m y length coordinate, m α absorptivity; also thermal diffusivity, m<sup>2</sup>/s

```
integrated band intensity, cm^{-1}/(g/m^2)
\alpha_{\mathbf{k}}
            molecular thermal diffusivity, m<sup>2</sup>/s
            line-width-to-spacing parameter; also dimension-
β
            less temperature ratio
            line half-width, cm<sup>-1</sup>; also Euler's constant
Υ
            0.5772156...
            thickness or half thickness, m; also vibrational
δ
            quantum number change
            emissivity
3
            heat exchanger effectiveness
<sup>€</sup>eff
            eddy diffusivity for heat
\epsilon_{H}
            eddy diffusivity for momentum
\varepsilon_{\rm M}
            line-width-to-spacing parameter; also fin effec-
η
            tiveness
θ
            polar angle
            mass absorption coefficient, (kg/m^2)^{-1}
K
            wavelength, µm
λ
            \cos \theta
μ
            magnetic permeability
            wavenumber, cm^{-1}; also kinematic viscosity, m^2/s
            frequency, s<sup>-1</sup>
νf
            molecular kinematic viscosity, m<sup>2</sup>/s
            3.1415926...
π
            reflectivity; also density kg/m<sup>3</sup>
P
            absorber partial density, kg/m<sup>3</sup>
\rho_{\mathbf{a}}
            electric change density
P
            Stefan-Boltzmann constant, 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4
```

σ

electrical conductivity  $\sigma_{\mathbf{e}}$  $\sigma_{\mathbf{r}}$ surface roughness, µm τ transmissivity optical depth at band head  $\tau_{H,k}$  $\boldsymbol{\tau}_{\text{WL}}$ wall layer transmissivity azimuthal angle hot band line spacing parameter, dimensionless hot band intensity parameter, dimensionless exponential band width, cm<sup>-1</sup> ω albedo for single scatter  $\omega_{s}$ Ω solid angle, sr

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# 8 RADIATION ACTING WITH CONDUCTION OR CONVECTION

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| 0./~ | Combined | pricriorita | 020 |

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# 1 INTRODUCTION

1.A Radiation Heat Transfer in Thermal Design. When does one consider radiation heat transfer, and when does one not? One does not consider radiation inside of a fluid that is highly opaque to the source spectrum. In a fluid such as water, the radiation is merely a contributor to what we know as thermal conductivity. Similarly one does not consider radiation inside a fluid that is perfectly transparent to the source spectrum. If there is no physical mechanism by which the fluid can absorb energy from radiation passing through it, then it follows from thermodynamics that it cannot emit radiation either, and it cannot be either heated or cooled by radiation. Such a fluid is said to be diathermanous. The walls surrounding such a fluid, however, may exchange heat radiation, but only if they are not isothermal. Thus one does not ordinarily consider radiation within the passages of a heat exchanger containing oil, water, or air. The first two are opaque. The last is diathermanous.

When two walls at different temperatures are in view of one another or one wall is in view of a participating medium (one neither opaque nor diathermanous) the radiation heat flux (W/m²) tends to be high when  $\Delta\sigma T^4$  is high, where  $\sigma$  is the Stefan-Boltzmann constant  $5.6697x10^{-8}$  W/m² K⁴. When  $\Delta T$  is small compared to the absolute temperature level,  $\Delta\sigma T^4$  can be written  $4\sigma T_m^3 \Delta T$  where  $T_m$  is the mean temperature level.

At 300 K the value for  $4\sigma T_m^{-3}$  is slightly over 6 W/m<sup>2</sup> K (circa 1 Btu/hr ft<sup>2</sup> R) on the same order as a natural convection heat transfer coefficient. At  $T_m = 2000$  K the value is nearly 300 times greater. From such a value, 1800 W/m<sup>2</sup> K, one can see why radiation contributes to film boiling heat transfer. Radiation is important when temperatures are high, distances are large (because convective heat transfer coefficients go like passage size D as D<sup>-1/5</sup> for turbulent flow or D<sup>-1</sup> for laminar flow), or under vacuum conditions when convective heat transfer coefficients are low because of the low fluid density.

1.B Thermodynamic Surfaces and Surface Systems. The thermal designer needs to know surface heat fluxes adjacent to the interface between phases. When one phase is highly opaque, and the other is not, the opaque surface system concept is used. Figure 1-1 depicts a surface system. The s-surface lies just outside the highly opaque phase; the u-surface lies just within it. The m-surface lies sufficiently below the phase interface so that (1) no radiation crossing the s and u surfaces is transmitted to the m-surface, and (2) the radiation flux crossing the m-surface is given by the radiation-diffusion equation and is included with the conduction. For no flow through the surfaces and negligible transient heat storage in the mass between the m and u surfaces, one has

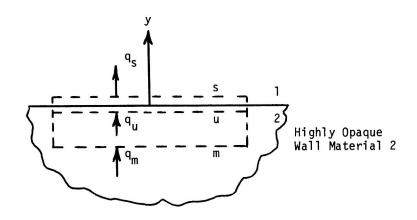


FIGURE 1-1

The s, u, and m Thermodynamic Boundaries for a Surface System