

牛津大学 研究生教材系列

Modern Classical Optics

现代经典光学

G. Brooker



科学出版社
www.sciencep.com

043/Y24

2009.

牛津大学研究生教材系列

Modern Classical Optics

现代经典光学

G. Brooker

科学出版社

北 京

图字:01-2007-2794 号

©Oxford University 2003

Modern Classical Optics was originally published in English in 2003. This adaptation is published by arrangement with Oxford University Press and is for sale in the Mainland(part) of the People's Republic of China only.

现代经典光学原书英文版于 2003 年出版。本改编版获得牛津大学出版社授权,仅限于在中华人民共和国大陆(部分)地区销售。

图书在版编目(CIP)数据

现代经典光学 = Modern Classical Optics: 英文/(英)布鲁克(Brooker, G.)著. —注释本. —北京:科学出版社,2009

(牛津大学研究生教材系列)

ISBN 978-7-03-023623-4

I. 现… II. 布… III. 光学-研究生-教材-英文 IV. 043

中国版本图书馆 CIP 数据核字(2008)第 201618 号

责任编辑:胡凯 王飞龙/责任印制:钱玉芬/封面设计:王浩

科学出版社 出版

北京东黄城根北街 16 号

邮政编码:100717

<http://www.sciencep.com>

新蕾印刷厂印刷

科学出版社发行 各地新华书店经销

*

2009 年 1 月第 一 版 开本:787×1092 1/16

2009 年 1 月第一次印刷 印张:26

印数:1~3 000 字数:501 000

定价:78.00 元

(如有印装质量问题,我社负责调换<长虹>)

Preface

The level of treatment in this book is that of the fourth year of an M. Phys. undergraduate course in the UK. However, I have tried to give descriptions that are simple enough to be followed by someone at an earlier stage who seeks a ‘different’ account of the more basic material. And graduates may find that some ‘well-understood’ ideas offer unexpected challenges. The topics included here are more than could be covered in the time available in any one undergraduate course, but different courses will, quite properly, make different selections of material.

I concentrate on ‘physical’ optics (light as a wave), and describe only as much geometrical optics as is really necessary. A thick lens is mentioned only three times. Lens design and optical aberrations are hardly mentioned at all, and then in terms of an optical transfer function rather than Seidel sums. I justify this exclusion on the ground that lens design is now wholly done by computer-aided optimization, description of which would require a very different style of presentation.

This book might better have been called ‘semi-classical optics’, since the photon nature of light is not ignored. Indeed, photon emission and detection are inherently quantum-mechanical. However, our main concern, the passage of light between emission and detection, can usually be treated classically. Those phenomena, such as entanglement or anti-bunching, that require ‘quantum optics’ proper lie outside our remit. Even so, I have tried, in Chapter 10, to explain where the interface lies between the (semi-)classical and quantum regimes.

In a book of this length, some selection of topics is unavoidable, even within physical optics. In particular I regret the omission of interference microscopes (too large a digression) and of adaptive optics applied to Earth-bound astronomical telescopes (too computational).

A book is a linear structure: from beginning to end. Understanding is not like that. It’s achieved by reading interactively: checking calculations; cross-linking new information with old; asking ‘what if’; thinking of implications and possible objections. Why is μ_r always assumed to be 1 at optical frequencies? Why is an electromagnetic wave always discussed in terms of its $\mathbf{E}$ -field when $\mathbf{B}$ is equally significant in the Maxwell equations? A Fabry–Perot and a thin film are very similar structures; why then are the methods of analysis so different? Can we trust the Kirchhoff-assumption boundary conditions used in diffraction, and how could we find out? Why are the fields inside a laser cavity mathematically similar to the wave functions for a simple harmonic oscillator? The bright student will want answers to such questions, and if a full answer

can't be given at this level then at least s/he wants to be given reassurance that the questions are not regarded as troublemaking. The less imaginative should be encouraged by example to see that these are the kind of questions that ought to be asked.

Critical questions shouldn't be allowed to clutter the exegesis, yet a place must be found for them. I have therefore put much exploratory material into 'problems'. A 'problem' is not necessarily an exercise set (though it may be) by the author or at his suggestion by a teacher; it is often an opportunity to make statements (to be checked) about special cases, to point out a link to a similar idea elsewhere, or to encourage discrimination by challenging the 'well-known'. In short, problems are meant to be read, as a commentary integral to the material; the reader can decide which ought to be worked through in detail as well. Most problems are referenced at an appropriate point in the text, to indicate that additional information or critical discussion is available. Solutions are given where mathematical steps are not all obvious or where the physics needs (even more!) discussion.

Problems are graded on a three-point scale from the most basic (a), usually easy, to the most challenging (c). A few 'entertainment' problems, labelled (e), are opportunities to take a wider view. Whatever its classification, each problem is intended to draw attention to some insight, technique or order of magnitude.

This book has grown out of lecture courses that I have given, over a number of years and at different levels. (The author sometimes complains that the Oxford Physics Department has type-cast him as an optician.) The content and attitude have been strongly influenced by frequent interaction with students, one or two at a time, in tutorial discussions that range over most areas of degree-level physics. To those students who asked searching questions or insisted on clearer explanations, I owe a great debt.

Several colleagues have contributed to this book, not always knowingly, by reading it in draft or via a number of discussions. I am particularly indebted to Professor D. N. Stacey and Professor C. E. Webb for their patience in the discussion of a number of tricky points. Dr S. J. Blundell is responsible for an improvement in the clarity of Chapter 5. Mr C. Goodwin very kindly supplied the practical detail on thin-film deposition in Chapter 6. Dr J. Halliday was most helpful in providing information on CD and DVD technology. Dr R. A. Taylor gave valuable help with the description of LEDs in Chapter 11.

I thank Springer-Verlag for permission to reproduce Figs 3.8, 3.9 and 3.12 from the *Atlas of Optical Phenomena*. I thank the IEEE and Dr H. Kogelnik for permission to reproduce Fig. 8.3. I thank Lucent Technology and Drs Fox, Li, Boyd and Gordon for permission to base Fig. 8.7 on a diagram that originally appeared in the *Bell System Technical Journal*. All other diagrams were prepared by the author.

Finally, I thank Marlene, who in her last three years shared less time with me than we both would have wished because this book was in preparation.

目 录

1 电磁学和光学基础	1
1.1 导引	1
1.2 Maxwell 方程组	1
1.3 线性各向同性介质	2
1.4 平面电磁波	3
1.5 能量流	4
1.6 标量波振幅	5
1.7 色散介质	6
1.8 输电线	6
1.9 基础(光线)光学	8
1.9.1 薄透镜	8
1.9.2 符号法则	10
1.9.3 球面折射	10
1.9.4 厚透镜	12
1.10 光线和波	12
问题	13
2 Fourier 级数和 Fourier 变换	21
2.1 导引	21
2.2 Fourier 级数:周期波谱	21
2.3 Fourier 级数:数学变形	22
2.4 Fourier 变换:非周期波谱	23
2.5 解析信号	25
2.6 δ 函数	26
2.7 频率和角频率	26
2.8 功率谱	27
2.9 Fourier 变换实例	27
2.9.1 单矩形脉冲	27
2.9.2 双矩形脉冲	29
2.9.3 δ 函数脉冲	29
2.9.4 规则的 δ 函数脉冲序列	30
2.9.5 随机的 δ 函数脉冲序列	30
2.9.6 无限正弦波	30
2.10 卷积和卷积定理	31

2.11	卷积的例子	31
2.12	Fourier 变换的信号选择	32
	问题	33
3	衍射	44
3.1	导引	44
3.2	单色球面波	44
3.3	Kirchhoff 衍射积分	44
3.4	Kirchhoff 边界条件	46
3.5	Kirchhoff 积分的化简	46
3.6	互补屏;Babinet 原理	49
3.7	Fraunhofer 条件 I	51
3.8	一维 Fraunhofer 衍射	51
3.9	二维 Fraunhofer 衍射	53
3.10	观测衍射的两种方式	55
3.11	Fraunhofer 衍射的实例	57
3.12	Fraunhofer 衍射和 Fourier 变换	61
3.13	Fraunhofer 条件 II;Rayleigh 距离和 Fresnel 数	62
3.14	Fraunhofer 条件 III;物和像	64
3.15	Fresnel 衍射	66
3.16	Fraunhofer 衍射和光学分辨率	68
3.17	Fourier 变换相关的光场面	69
3.18	Fraunhofer 边界条件	70
	问题	72
4	衍射光栅	86
4.1	导引	86
4.2	透射光栅	86
4.3	多元图样	86
4.4	反射光栅	88
4.5	闪耀	89
4.6	光栅光谱仪	90
4.7	光谱分辨率	92
4.8	制作光栅	94
4.9	一些诀窍	95
4.9.1	标准光谱	95
4.9.2	光源校正	96
4.9.3	用摄谱仪缩短曝光时间	97
4.9.4	真空设备	97
4.9.5	双单色仪	97

4.9.6 发明者的天堂	98
4.10 超越简单理论	99
问题	100
5 Fabry-Perot 干涉仪	109
5.1 导引	109
5.2 基本理论	109
5.3 基本装置	111
5.4 细度的意义	112
5.5 自由光谱范围和分辨率	113
5.5.1 自由光谱范围	113
5.5.2 分辨率	114
5.6 标准条纹图样分析	115
5.7 Fabry-Perot 板的平面度和平行度	117
5.8 设计 Fabry-Perot 干涉仪	118
5.9 用 Fabry-Perot 干涉仪进行光谱分析	120
5.10 Fabry-Perot 干涉仪的延伸	122
问题	122
6 薄膜	132
6.1 导引	132
6.2 单层薄膜的简单计算	132
6.3 矩阵“中”场振幅消去法	134
6.4 反射波和透射波	135
6.5 光阻的概念	135
6.6 高反射率面镜	137
6.7 无反镀层	139
6.8 干涉滤波器	140
6.9 薄膜沉积	141
问题	144
7 光束矩阵和 Gauss 光束	150
7.1 导引	150
7.2 光线光学中的矩阵方法	150
7.3 转换矩阵和折射矩阵	151
7.4 反射	152
7.5 球面波	153
7.6 Gauss 光束	154
7.7 Gauss 光束的性质	155
7.8 符号规则	156
7.9 Gauss 光束的传播	156

7.10	电场和磁场	159
问题		160
8	光腔	170
8.1	导引	170
8.2	Gauss-Hermite 光束	170
8.3	腔共振	171
8.4	腔模	172
8.5	低损耗模的条件	174
8.6	腔模的形状	177
8.7	纵模	178
8.8	高损耗腔	178
8.9	对称共焦腔	180
8.10	共焦 Fabry-Perot 干涉仪	181
8.11	激光腔几何形状的选择	182
8.12	反模的选择	182
8.13	模匹配	183
问题		183
9	干涉:定性描述	194
9.1	导引	194
9.2	术语	194
9.3	杨氏条纹:频率范围的容偏	196
9.4	杨氏条纹:准直容偏	196
9.5	干涉区域	198
9.6	Michelson 恒星干涉仪	199
9.7	孔径综合法	200
9.8	纵向相干和横向相干	201
9.9	两列平行平面波的干涉	202
9.10	快、慢探测器	204
9.11	相干时间和相干长度	204
9.12	用 Michelson 干涉仪研究长缝干涉	207
9.13	条纹可见度	209
9.14	量级	210
9.15	讨论	210
9.15.1	激光的干涉	211
9.15.2	杨氏双缝	211
9.15.3	快、慢探测器	212
9.15.4	光栅单色仪	212
9.15.5	极化和非极化光	213

问题	213
10 干涉:相关函数	219
10.1 导引	219
10.2 相关函数的定义	219
10.3 白相关和 Mechelson 干涉仪	221
10.4 归一化白相关函数	221
10.5 条纹可见度	222
10.6 Wiener-Khintchine 定理	222
10.7 Fourier 变换谱	223
10.8 部分相干	226
10.9 van Cittert-Zernike 定理	227
10.10 密度相关	229
10.11 混沌光和激光	229
10.12 Hanbury Brown-Twiss 实验	230
10.13 用强度关联测量恒星直径	232
10.14 经典光学和量子光学	233
问题	235
11 光学实践:光束扩展、干涉计量、条纹定域	249
11.1 光束扩展	249
11.2 能流:光束扩展和辐照亮度	249
11.3 光束扩展守恒和辐照亮度守恒	250
11.4 纵模和横模	251
11.5 光束扩展域和相干域	251
11.6 场模和熵	252
11.7 几类光源的照射	253
11.7.1 黑体辐射	253
11.7.2 气体放电照射	253
11.7.3 发光二极管照射	254
11.8 光束扩展和干涉仪	254
11.9 光束扩展和光谱仪	255
11.10 Fourier 变换光谱仪的设计研究	256
11.11 条纹定域	256
问题	260
12 成像:衍射理论	272
12.1 导引	272
12.2 横向相干照明成像	272
12.3 想光学系统成像	274
12.4 非理想光学系统成像	275

12.5	显微镜分辨率:Abbe 理论	275
12.5.1	Abbe 理论导引	276
12.5.2	Abbe 理论解读	276
12.6	改善显微镜	278
12.7	相称	279
12.8	暗场照明	281
12.9	纹影	281
12.10	变迹	282
12.11	全息法	283
12.12	点扩散函数	283
12.13	光学传输函数、调制传输函数	284
	问题	286
13	全息法	292
13.1	导引	292
13.2	特例:平面波物光束和平面波参考光束	293
13.3	参考光束的强度	294
13.4	感光乳剂的响应	295
13.5	全息理论	297
13.6	成像	299
13.7	把全息图一分为二会怎样?	300
13.8	改变光学几何关系二次成像	301
13.9	厚感光乳剂的影响	302
13.10	相位全息图	303
13.11	Gabor 全息图	303
13.12	实践	304
13.13	全息的应用	305
	问题	306
14	光纤	311
14.1	导引	311
14.2	光纤光学基础	312
14.3	横模	313
14.4	色散	316
14.4.1	材料色散	317
14.4.2	模间和模内色散	317
14.5	多模光纤	318
14.6	单模光纤	318
	问题	319

15 极化	324
15.1 导引	324
15.2 各向异性介质	324
15.3 各向同性的数学	324
15.4 张量 ϵ_{ij} 的理解	325
15.5 Farady 效应	326
15.6 光活性	327
问题	327
16 两种现代光学仪器	337
16.1 导引	337
16.2 光盘的描述	337
16.3 光盘的编码机制	338
16.4 读盘的光学原理	339
16.5 反馈系统	340
16.5.1 光道校正	340
16.5.2 聚焦校正	341
16.6 CD-ROM	342
16.7 DVD	342
16.8 共焦显微镜	343
16.9 共焦显微镜的分辨率	344
16.10 共焦显微镜的焦深	344
问题	345
问题注解	346
参考文献	383
索引	387

Contents

1	Electromagnetism and basic optics	1
1.1	Introduction	1
1.2	The Maxwell equations	1
1.3	Linear isotropic media	2
1.4	Plane electromagnetic waves	3
1.5	Energy flow	4
1.6	Scalar wave amplitudes	5
1.7	Dispersive media	6
1.8	Electrical transmission lines	6
1.9	Elementary (ray) optics	8
1.9.1	The thin lens	8
1.9.2	Sign conventions	10
1.9.3	Refraction at a spherical surface	10
1.9.4	The thick lens	12
1.10	Rays and waves	12
	Problems	13
2	Fourier series and Fourier transforms	21
2.1	Introduction	21
2.2	Fourier series: spectrum of a periodic waveform	21
2.3	Fourier series: a mathematical reshape	22
2.4	The Fourier transform: spectrum of a non-periodic waveform	23
2.5	The analytic signal	25
2.6	The Dirac δ -function	26
2.7	Frequency and angular frequency	26
2.8	The power spectrum	27
2.9	Examples of Fourier transforms	27
2.9.1	A single rectangular pulse	27
2.9.2	The double pulse	29
2.9.3	A δ -function pulse	29
2.9.4	A regular array of δ -functions	30
2.9.5	A random array of δ -functions	30
2.9.6	An infinite sinewave	30
2.10	Convolution and the convolution theorem	31
2.11	Examples of convolution	31
2.12	Sign choices with Fourier transforms	32
	Problems	33

3	Diffraction	44
3.1	Introduction	44
3.2	Monochromatic spherical wave	44
3.3	The Kirchhoff diffraction integral	44
3.4	The Kirchhoff boundary conditions	46
3.5	Simplifying the Kirchhoff integral	46
3.6	Complementary screens: the Babinet principle	49
3.7	The Fraunhofer condition I: provisional	51
3.8	Fraunhofer diffraction in ‘one dimension’	51
3.9	Fraunhofer diffraction in ‘two dimensions’	53
3.10	Two ways of looking at diffraction	55
3.11	Examples of Fraunhofer diffraction	57
3.12	Fraunhofer diffraction and Fourier transforms	61
3.13	The Fraunhofer condition II: Rayleigh distance and Fresnel number	62
3.14	The Fraunhofer condition III: object and image	64
3.15	The Fresnel case of diffraction	66
3.16	Fraunhofer diffraction and optical resolution	68
3.17	Surfaces whose fields are related by a Fourier transform	69
3.18	Kirchhoff boundary conditions: a harder look	70
	Problems	72
4	Diffraction gratings	86
4.1	Introduction	86
4.2	A basic transmission grating	86
4.3	The multiple-element pattern	86
4.4	Reflection grating	88
4.5	Blazing	89
4.6	Grating spectrometric instruments	90
4.7	Spectroscopic resolution	92
4.8	Making gratings	94
4.9	Tricks of the trade	95
4.9.1	Normal spectrum	95
4.9.2	Correct illumination	96
4.9.3	Shortening exposure times with a spectrograph	97
4.9.4	Vacuum instruments	97
4.9.5	Double monochromator	97
4.9.6	An inventor’s paradise	98
4.10	Beyond the simple theory	99
	Problems	100
5	The Fabry–Perot	109
5.1	Introduction	109
5.2	Elementary theory	109
5.3	Basic apparatus	111
5.4	The meaning of finesse	112
5.5	Free spectral range and resolution	113
5.5.1	Free spectral range	113

5.5.2	Resolution	114
5.6	Analysis of an étalon fringe pattern	115
5.7	Flatness and parallelism of Fabry–Perot plates	117
5.8	Designing a Fabry–Perot to do a job	118
5.9	Practicalities of spectroscopy using a Fabry–Perot	120
5.10	The Fabry–Perot as a source of ideas	122
	Problems	122
6	Thin films	132
6.1	Introduction	132
6.2	Basic calculation for one layer	132
6.3	Matrix elimination of ‘middle’ amplitudes	134
6.4	Reflected and transmitted waves	135
6.5	Impedance concepts	135
6.6	High-reflectivity mirrors	137
6.7	Anti-reflection coatings	139
6.8	Interference filters	140
6.9	Practicalities of thin-film deposition	141
	Problems	144
7	Ray matrices and Gaussian beams	150
7.1	Introduction	150
7.2	Matrix methods in ray optics	150
7.3	Matrices for translation and refraction	151
7.4	Reflections	152
7.5	Spherical waves	153
7.6	Gaussian beams	154
7.7	Properties of a Gaussian beam	155
7.8	Sign conventions	156
7.9	Propagation of a Gaussian beam	156
7.10	Electric and magnetic fields	159
	Problems	160
8	Optical cavities	170
8.1	Introduction	170
8.2	Gauss–Hermite beams	170
8.3	Cavity resonator	171
8.4	Cavity modes	172
8.5	The condition for a low-loss mode	174
8.6	Finding the mode shape for a cavity	177
8.7	Longitudinal modes	178
8.8	High-loss cavities	178
8.9	The symmetrical confocal cavity	180
8.10	The confocal Fabry–Perot	181
8.11	Choice of cavity geometry for a laser	182
8.12	Selection of a desired transverse mode	182
8.13	Mode matching	183
	Problems	183

9 Coherence: qualitative	194
9.1 Introduction	194
9.2 Terminology	194
9.3 Young fringes: tolerance to frequency range	196
9.4 Young fringes: tolerance to collimation	196
9.5 Coherence area	198
9.6 The Michelson stellar interferometer	199
9.7 Aperture synthesis	200
9.8 Longitudinal and transverse coherence	201
9.9 Interference of two parallel plane waves	202
9.10 Fast and slow detectors	204
9.11 Coherence time and coherence length	204
9.12 A Michelson interferometer investigating longitudinal coherence	207
9.13 Fringe visibility	209
9.14 Orders of magnitude	210
9.15 Discussion	210
9.15.1 What of lasers?	211
9.15.2 The Young slits: another look	211
9.15.3 Fast and slow detectors: another look	212
9.15.4 Grating monochromator: another look	212
9.15.5 Polarized and unpolarized light	213
Problems	213
10 Coherence: correlation functions	219
10.1 Introduction	219
10.2 Correlation function: definition	219
10.3 Autocorrelation and the Michelson interferometer	221
10.4 Normalized autocorrelation function	221
10.5 Fringe visibility	222
10.6 The Wiener–Khintchine theorem	222
10.7 Fourier transform spectroscopy	223
10.8 Partial coherence: transverse	226
10.9 The van Cittert–Zernike theorem	227
10.10 Intensity correlation	229
10.11 Chaotic light and laser light	229
10.12 The Hanbury Brown–Twiss experiment	230
10.13 Stellar diameters measured by intensity correlation	232
10.14 Classical and quantum optics	233
Problems	235
11 Optical practicalities: étendue, interferometry, fringe localization	249
11.1 Introduction	249
11.2 Energy flow: étendue and radiance	249
11.3 Conservation of étendue and radiance	250
11.4 Longitudinal and transverse modes	251
11.5 Étendue and coherence area	251

11.6	Field modes and entropy	252
11.7	Radiance of some optical sources	253
11.7.1	Radiance of a black body	253
11.7.2	Radiance of a gas-discharge lamp	253
11.7.3	Radiance of a light-emitting diode (LED)	254
11.8	Étendue and interferometers	254
11.9	Étendue and spectrometers	255
11.10	A design study: a Fourier-transform spectrometer	256
11.11	Fringe localization	256
	Problems	260
12	Image formation: diffraction theory	272
12.1	Introduction	272
12.2	Image formation with transversely coherent illumination: informal	272
12.3	Image formation: ideal optical system	274
12.4	Image formation: imperfect optical system	275
12.5	Microscope resolution: Abbe theory	275
12.5.1	Abbe theory: introduction	276
12.5.2	Abbe theory: explanation	276
12.6	Improving the basic microscope	278
12.7	Phase contrast	279
12.8	Dark-ground illumination	281
12.9	Schlieren	281
12.10	Apodizing	282
12.11	Holography	283
12.12	The point spread function	283
12.13	Optical transfer function; modulation transfer function	284
	Problems	286
13	Holography	292
13.1	Introduction	292
13.2	Special case: plane-wave object beam and plane-wave reference beam	293
13.3	The intensity of the reference beam	294
13.4	The response of a photographic emulsion	295
13.5	The theory of holography	297
13.6	Formation of an image	299
13.7	What if we break a hologram in half?	300
13.8	Replay with changed optical geometry	301
13.9	The effect of a thick photographic emulsion	302
13.10	Phase holograms	303
13.11	Gabor's holograms	303
13.12	Practicalities	304
13.13	Applications of holography	305
	Problems	306