# Fundamentals of Astrodynamics and Applications

**Second Edition** 

# David A. Vallado

with contributions by Wayne D. McClain





# Fundamentals of Astrodynamics and Applications

### **Second Edition**

# David A. Vallado

with technical contributions by Wayne D. McClain



Published Jointly by

Microcosm Press

El Segundo, California



Kluwer Academic Publishers

Dordrecht / Boston / London

#### Library of Congress Cataloging-in-Publication Data

A C.I.P. Catalogue record for this book is available from the Library of Congress

ISBN 1-881883-12-4 (pb) (acid-free paper) ISBN 0-7923-6903-3 (hb) (acid-free paper)

Published jointly by
Microcosm Press
401 Coral Circle, El Segundo, CA 90245 USA
and
Kluwer Academic Publishers,
P.O. Box 17, 3300 AA Dordrecht, The Netherlands.

Kluwer Academic Publishers incorporates the publishing programmes of D. Reidel, Martinus Nijhoff, Dr. W. Junk and MTP Press.

Sold and distributed in the USA and €anada \*by Microcosm, Inc.
401 Coral Circle, El Segundo, CA 90245 USA and Kluwer Academic Publishers,
101 Philip Drive, Norwell, MA 02061 USA

In all other countries, sold and distributed by Kluwer Academic Publishers Group, P.O. Box 322, 3300 AA Dordrecht, The Netherlands.

Printed on acid-free paper

All Rights Reserved
© 2001 David A. Vallado

No part of the material protected by this copyright notice may be reproduced or utilized in any form or by any means, electronic or mechanical, including photocopying, recording, or by any information storage and retrieval system, without written permission from the publishers.

Printed in the United States of America

#### Foreword

At Microcosm, when an astrodynamics problem arises it has become standard practice to "look it up in Vallado." In most cases, the answer is there, and in a form that can be readily used and applied to the spaceflight problem at hand. I believe it is for this reason that *Fundamentals of Astrodynamics and Applications* has so rapidly become the standard astrodynamics reference for those of us in the business of spaceflight.

Thus, it is truly a pleasure to welcome the revised and expanded second edition of "Vallado" to the Space Technology Library. While the first edition was an exceptionally useful and popular book throughout the community, there are a number of reasons why we believe the second edition will be even more so. The text has expanded to 958 pages. There are many reworked examples and derivations. Entirely new topics include ground illumination calculations, Moon rise and set, and a listing of relevant InterNet sites. There is an improved and expanded discussion of coordinate systems, orbit determination, and differential correction. Perhaps most important is that all of the software routines described in the book are now available for free in FORTRAN, PASCAL, and C on the book website, www.smad.com/Vallado. With extensive references to both specific equations and the section of the book where the topic is discussed, this software is the perfect practical companion to Vallado's book. This makes the second edition even more valuable as both a text and reference.

We hope that you will find this book to be as useful and productive as those of us at Microcosm have. To me, a good measure of its real utility is that nearly every engineer in our Space Systems Division has spent their own hard-earned money and bought a copy for themselves. The book is not only an excellent text, it is a superb reference that will be used over and over in practical astrodynamics.

To you, our readers, we are certain that you will enjoy the clarity and simplicity with which Dave both derives and explains modern, applied astrodynamics. To Dave Vallado, Welcome Aboard! We are delighted to have you join the Space Technology Library.

James R. Wertz El Segundo, CA February 2001 jwertz@smad.com

#### Preface to the Second Edition

My goal in writing this book is to augment and consolidate some of the information other astrodynamics books contain and, perhaps more importantly, to present and describe the computational fundamentals of astrodynamics. During the early 1960s, tremendous progress took place in astrodynamics, in part due to the U.S.'s drive to reach the Moon and to stay technologically superior during the Cold War. While the theories advanced by large steps, the computer code took only a "small step" because of available computing languages and hardware. But in the 1970s and 1980s, the computer industry took gigantic steps, so now is the time to capitalize on advancements in both areas.

You're probably wondering, does this book present anything new? Well, yes and no. The dynamics have been in effect since the beginning of time, and most of our mathematics hasn't changed much since its discovery over 300 years ago. Today, advances in computer technology allow students and practitioners to determine orbits with high precision on a PC, but they need a compendium of *well-documented* algorithms to use that power. Hence, this book.

The days of studying just one discipline are nearly over. Today's scientists should study computer science in *addition* to rigorous engineering principles—after all, that's how we do our business. As older managers retire and newer computer systems arrive, the need is more pressing than ever for a new look at *traditional* astrodynamics. The days of "well, it works" are long gone.

The keys for the future are standardizing and documenting, modularity, efficiency, and maintainability. This book attempts to use the new programming capabilities to exploit each of these areas. Standardizing and documenting can't be overemphasized. I've witnessed tremendous efforts in astrodynamics and, on every project, I always came away wishing that someone had written a few more steps, or had provided a test case to check how the program worked. Too often, books are written mainly for academics; writers forget that engineers need simple algorithms and examples to use in everyday operations. Operational "documentation" is often severely limited or nonexistent. Although I use standard notation as much as possible and work hard to match existing practices, it's impossible to be consistent with all sources.

I've included many derivations because I've seen too many engineers spend excessive time trying to recreate something that is "easy to verify." Recall the infamous scribbling of Pierre de Fermat (1601-1665) in the margin of a page where he omitted the proof because surely everyone could understand it and there wasn't enough room to say more (*Scientific American*, Oct. 1978, Oct. 1993, Oct. 1994, and Nov. 1997, 68-73). Fermat's last theorem remained a mystery for 350 years and though proven by Andrew Wiles in 1994, the lack of Fermat's written derivation generated controversy for several hundred years! Redoing fundamental problems is useful to learn the basics, but we can save steps (and therefore time and money) by presenting the material straightforwardly, with consistent notation.

Modularity is closely aligned with standardizing and documenting, but it's distinct. I've become accustomed to the old axiom "this program already does everything you

need," only to find that your application requires just 5% of the code, and you can't pull out the routines you need because all routines are interconnected. In this book I present many independent algorithms. Admittedly, some routines will share algorithms, but with a little thought, this becomes only a minor issue.

Efficiency is important, yet elusive. During the development, we can consider some simple guidelines that will make routines run much faster. I've given you lots of examples and hints throughout the book.

Finally, at times it will seem as though I'm leading you by the hand through a problem, and at others, you'll need to scratch some intermediate equations in the margin (make sure you do record your notes!). My goal in writing this book has been (within some reasonable bounds for the overall length) to derive and detail each algorithm to about the same level to provide a solid foundation. Because individual abilities vary widely, you may find routines you've worked with to be too simple and unknown algorithms to be confusing. I've tried to minimize the confusion, and I'll appreciate your comments for improvement.

You may need some prerequisites for this book. It tries to bridge topics in astrodynamics from undergraduate courses with follow-on, graduate-level topics. It's intended for upper-level undergraduate classes in astrodynamics, as well as for some graduate-level classes. From my experience in laboratory research, operational environments, and instruction, I've tried to present the material with a practical flavor. The book covers theory, applications, and interpretations through numerous examples. The idea is to blend real-world operations with a theoretical understanding of the mechanics at work.

I've coded almost all the algorithms to verify each technique. Actual computer source code is available via the Internet. It details test cases and code in PASCAL, FORTRAN, and C. These languages should permit relatively easy conversion to other languages.

This book is intended to introduce astrodynamics to the undergraduate, and to lay a solid foundation for graduate-level studies. As such, it covers a wide scope of topics, some of which are primarily for academic purposes, while others are for general interest. Topics in astronomy are introduced where relevant to astrodynamics and the study of Earth-orbiting satellites, as well as for general information. Having a degree or studying astrodynamics, you'll invariably get questions about "what star is that?" This text probably will not prepare you for that task. However, you can consult any number of introductory astronomy books to provide an answer to that question!

The first 7 chapters relate primarily to two-body motion. In particular, the first five chapters have been re-ordered from the first edition to better match undergraduate curriculums. The text now begins with the two-body equations of motion, and Kepler's solutions. This permits initial analysis of orbits, and sets the stage for the coordinate and time system discussions in Chap. 3. The old Chap. 4 was broken into two chapters and combined with information form initial orbit determination and estimation. This combination provides a more consolidated discussion of observations and celestial phenomena.

The last 4 chapters discuss more advanced topics, and concentrate on propagation (numerical and analytical), and differential correction. These topics form the heart of orbit determination systems, whether they handle all 10,000 satellites in Earth orbit, or a single satellite. My placement of the material in this order isn't intended to restrict first

or second courses from using these chapters. In fact, I'd expect a first course to incorporate topics such as predicting a satellite's location over a ground site (Chap. 11), or an introductory numerical propagation exercise (Chap. 8), or an initial orbit determination method (Gibbs or Herrick-Gibbs Chap. 7). I tired to keep all the similar material in the same chapter to avoid looking back several chapters to see the various approaches to accomplish a task.

I show many detailed derivations, culminating in an algorithm, suitable for programming a computer. I provide comments about computer programming because of the importance computers have become today in all disciplines. For the seasoned veteran, the algorithm may be trivial, and the computer coding comments may be unwarranted. For the student, these extra items may be the means to understanding, if not simply getting the program to work! The derivation ensures the assumptions, methodology, and rationale are clear for each technique.

For academic courses, several prerequisites are necessary. I have included a very brief review of vectors and matrices, and some numerical techniques in Appendix C. These are not intended to fully explain the concepts, but rather to introduce my notation and nomenclature.

For the casual reader, instructor, etc. who wishes further information in certain areas, I've listed many references in the reference section. I include the number of the chapter they are cited in, and the references indicate page numbers to help you find the citation quickly. I recommend a few texts that are especially good in certain areas. This list is not exhaustive!

Advanced Astrodynamics: Battin (1987), Szebehely (1967), Stiefle and Scheifele (1971)

Analytical propagation: Brouwer and Clemence (1961)

Attitude determination: Wertz (1978)

Estimation: Gelb (1989), Bierman (1977)
General Astrodynamics: Escobal (1985), Danby (1992)
Gravity, geodesy: Lambeck (1988), Kaula (1966)

Mathematics: Kreyzig (1983), Numerical Recipes (1992), Sadler (1974)

Time, coordinates, astronomy: Seidelmann (1992), Green (1988)

For academic courses, several pre-requisites are necessary. First, an understanding of analytical geometry, introductory calculus, differential equations, and basic computer programming are very useful. I mention several comments that apply to any computer programming language, and basic skills will be very useful. Data structures can also be helpful, depending on the additional computer programs that are assigned during a class. The estimation sections will benefit from introductory knowledge of statistics and probability theory. Although I explain several concepts, a single text can't explain everything!

Finally, although I've made every attempt to correct mistakes in the book, I'm sure there will be some errors. I encourage you to contact myself, or Wayne McClain by e-mail for any questions you discover. *All* comments are important!

valladodl@worldnet.att.net or wmcclain@draper.com

Questions regarding the publication of the book should be sent to Microcosm Press at bookproject@smad.com

# Acknowledgments

I must credit the numerous people who made this project a success. First and foremost, I thank my parents who have encouraged me throughout my life. My early childhood memories are of the 1960s, when the Cold War raged, the Cuban missile crisis arose, and Vietnam erupted. To their credit, my parents somehow diverted my attention toward one of the most important events to occur in recent history: the effort to land a man on the Moon. I'll never forget that feeling I had, glued to the TV (VCRs and DVDs weren't available yet) as Neil Armstrong walked the surface of the Moon. That event, coupled with my father's example as an engineer on the Manhattan Project during World War II, compelled me to become an engineer. Space was the obvious choice as the final frontier to explore.

Any large work needs solid technical reviews, but I'm especially thankful for those of Wayne McClain. As the primary technical reviewer, he helped me refine this book into a polished presentation. His insight and extensive experience were irreplaceable. Also, his patience with my ever-changing ideas and endless questions was truly remarkable! He provided extensive technical contributions to the book resulting from his experience in the field. The Charles Stark Draper Laboratory also deserves credit for granting Wayne McClain the time necessary to provide the thorough review of this work.

I'd also like to credit the people who allowed me to adapt parts of their work. I've listed each person alphabetically to avoid any favoritism—every comment was useful!

- Salvatore Alfano provided many papers for this book on continuous-thrust modeling (with Jim Thorne), close approach (with David Negron), rise/set calculations, and the method of ratios (also with David Negron). These works add a lot of practical insight and applications to my presentations.
- Leon Blitzer graciously permitted use of his *Handbook For Orbital Perturbations* (1970) for the perturbations chapters. Although I changed the look of his handbook material, much of the content remains, and I tried to keep the original intent he achieved almost thirty years ago.
- Carole Boelitz permitted me to include sections from her thesis in the perturbations sections. She also provided valuable reviews and derivations.
- Chapter 10 was initially based on a handout Jack Ferguson developed at the U.S. Air Force Academy in 1978. Jack Ferguson, others, and I reworked this handout in 1992 and I've expanded it again for this application.
- Daniel Fonte also allowed me to use sections of his thesis and his recent technical
  papers to give great credibility to the perturbations section. This area has been
  neglected for much too long.
- Tom Logsdon and Warner Miller provided initial thoughts and insight for the GPS section for Chap. 11. Mike Gabor helped refine this chapter into its final form.
- Jim Thorne allowed me to use sections from his continuous thrust and series reversion for solution of the Lambert problem. Some examples are from his work.

• Vasiliy Yurasov provided details on the Russian Space Surveillance System, and their atmospheric models. He also submitted an update to the GOST model, which is contained in Appendix B. This represents a unique collection of data which was formerly difficult to obtain.

I received tremendous support from many people throughout the astrodynamics community. In particular,

 Gerry Baumgartner and Michael Gabor provided a very detailed examination of the complete text. They gave extensive derivations and provided a level of confidence in many of the difficult sections.

Other people worked entirely through the text, some more than once:

• Don Danielson (and many of his students), Chuck Fosha (and his students), Kim Luu, Richard Rast, and Gary Yale.

Other individuals concentrated on certain chapters, relying on their expertise, which worked especially well considering the broad range of topics covered in the book:

 Carole Boelitz, Paul Cefola, Jack Ferguson, Jim Hall, Richard Hujsak, John Junkins, David Negron, Jr., Beny Neta, David Richardson, Ken Seidelmann, Wayne Smith, Sid Sridharan, and Jim Thorne.

Several people reviewed certain sections in detail:

 Cheryl Arney, Eric Beck, Eric Claunch, Steve Crumpton, Gim Der, Karl Jensen, Ron Lisowski, Anne Long, Ron Madler, Warner Miller, Steve Nerem, Dave Pohlen, Chris Sabol, David Spencer, David Suzuki, Nate Titus, Mark Tollefson, Scott Wallace, Kitt Carlton-Wippern, Gary Wirsig, and Lora Worth.

Any book written mostly by an engineer needs fine tuning to become a complete work. For the first edition, Daryl Boden assisted as a technical editor and provided technical insight for the text. Perry Luckett, a former professor of mine, really came to the rescue. He clarified ideas through more direct language, standardized or sharpened figures and tables, and kept my English in line!

I enjoyed continued support from the leadership at Air Force Research Laboratory, including Richard Davis, Eugene Dionne, Michael Heil, Henry Pugh, and Edward Duff. Their support for, and perseverance with, this long project was invaluable.

Several people formatted the first edition in FrameMaker Ver. 4.02. I began the typing and oversaw the operation throughout. Elsie Lugo deserves great credit for helping me create the first draft. Linda Pranke continued as production editor through the many intermediate drafts until the book was close to publication. She not only endured my cryptic notations but attended to every detail and caught inconsistencies that might have confused readers. She really made this book happen—despite my endless changes!

The first edition also owes many thanks to Margaret Hollander and Wiley Larson for getting the book published and on the streets.

Many people helped review the book by asking questions over the last few years. Their efforts were an important factor in finding errors and suggesting corrections. I am very grateful for their efforts and persistence to find me!

Salvatore Alfano, Jeff Back, Robert Bell, Beth Blakney, Daryl Boden, George Born, Scott Carter, Paul Cefola, Vince Chioma, Ching-pyang Chang, Shannon Coffey, Steve Crumpton (and his many proof-reading efforts!), Don Danielson (and his colleagues), Lee Dewald, Jack Fisher, Chuck Fosha, Robert Gay, Robert Good, Tom Jerardi, Paul Jermyn, Felix Hoots, Steven Hughes, Haisam Ido, Dave Kaslow, Clark Keith, Doug Kirkpatrick, Kim Luu, Benjamin Mains, Pat Marshall, Mark Matney, David Mikolajczak, Andrey Nazarenko, Steven Nerem, Michael Owen, Bahvesh Patel, Herb Reynolds, Chris Sabol, John Seago, Dale Shell, Sean Sherrard, Kirk Sorenson, David Spencer, Steve Stogsdill, Jim Thorne, Jim Torley, Rebecca Vallado, Jerry Vetter, Jason Yingst, Vasiliy Yurasov, and George Zabriskie.

I'd like to mention John Seago and Mike Gabor in particular as they helped refine the second edition with numerous comments and suggestions. Their insight, in-depth knowledge, and attention to detail were invaluable in making the second edition a reality.

I completed the second edition in FrameMaker Ver 5.5 which greatly improved the speed of processing the text. Finally, and certainly not least, Jim Wertz helped get this edition published. His counsel has been important throughout this effort, and I am very grateful for his assistance!

David A. Vallado
Denver, CO
February 2001
davallado@west.raytheon.com
valladodl@worldnet.att.net

# **Table of Contents**

Chapter 1	<b>Equations of Motion</b>		1
	1.1	History 1.1.1 Ancient Era 1.1.2 The Copernican Revolution 1.1.3 Kepler's Laws 1.1.4 Newton's Laws	I 1 7 9 10 12
	1.2	1.1.5 Other Early Astrodynamic Contributions Geometry of Conic Sections	12
		1.2.1 Basic Parameters	12
	1.3	Two-body Equation	20
		<ul> <li>1.3.1 Assumptions for the Two-Body Equation</li> <li>1.3.2 Specific Angular Momentum</li> <li>1.3.3 Specific Mechanical Energy</li> <li>1.3.4 Kepler's First Law (Trajectory Equation)</li> <li>1.3.5 Kepler's Second and Third Laws</li> <li>1.3.6 Velocity Formulas</li> </ul>	23 23 25 27 29 32
	1.4	Three-body and <i>n</i> -body Equations	33
		1.4.1 Inertial, Relative, and Barycentric Formulas	33
		1.4.2 Ten Known Integrals	37
		1.4.3 General Three-body Problem	40
Chapter 2	Kep	oler's Equation and Kepler's Problem	49
	2.1	Historical Background	49
	2.2	Kepler's Equation	51
		2.2.1 Alternate Formulation for Eccentric Anomaly	57
		2.2.2 Formulation for the Parabolic Anomaly	58
		2.2.3 Formulation for the Hyperbolic Anomaly 2.2.4 Universal Formulation	60 66
		2.2.5 Solutions of Kepler's Equation	71
		2.2.6 Summary and Related Formulas	85
	2.3	Kepler's Problem	87
		2.3.1 Solution Techniques	88
	2.4	Satellite State Representations	103
		2.4.1 Classical Orbital Elements (Keplerian)	104
		2.4.2 2-line Element Sets	113
		2.4.3 Equinoctial Elements	116 117
	2.5	2.4.4 Canonical Elements Application: Orbital Elements from <i>r</i> and <i>v</i>	118
	2.5	Application: r and v from Orbital Elements	122
	2.7	• •	126
	2.7	Application: Groundtracks Application: Find Time of Flight (FINDTOF)	130
	2.0	Application. This Time of Flight (Flight of )	150
Chapter 3	Coc	ordinate and Time Systems	135
	3.1	Historical Background	135
	3.2	The Earth	137
-		3.2.1 Location Parameters	139
		3.2.2 Shape of the Earth	140
		3.2.3 Gravitational Model	147

	3.3	Coordinate Systems	151
		3.3.1 Interplanetary Systems	156
		3.3.2 Earth-based Systems	157
		3.3.3 Satellite-based Systems	161
	3.4	Coordinate Transformations	165
		3.4.1 Coordinate Rotation	166
		3.4.2 Rotating Transformations	171
		3.4.3 Common Transformations 3.4.4 Application: Converting IJK To Latitude and Longitude	172 174
	3.5	3.4.4 Application: Converting IJK To Latitude and Longitude Time	179
	5.5	3.5.1 Solar Time and Universal Time	182
		3.5.2 Sidereal Time	189
		3.5.3 Atomic Time	193
		3.5.4 Dynamical Time	194
	3.6	Time Conversions	198
		3.6.1 DMS to Rad / Rad to DMS	199
		3.6.2 HMS to Rad / Rad to HMS	200
		3.6.3 HMS to Time of Day/ Time of Day to HMS	201 201
		3.6.4 YMD to Day of Year/ Day of Year to YMD 3.6.5 YMDHMS to Days / Days to YMDHMS	201
		3.6.6 Julian Date to Gregorian Date	203
	3.7	Transforming Celestial and Terrestrial Coordinates	205
	3.7	3.7.1 Non-rotating Origin	207
		3.7.2 FK5 Reduction	211
		3.7.3 Precession (FK5)	214
		3.7.4 Nutation (FK5)	216
		3.7.5 Sidereal Time (FK5)	219 219
		3.7.6 Polar Motion (FK5)	220
		3.7.7 Summary (FK5) 3.7.8 FK4 Reduction	227
	3.8	Earth Models and Constants	228
		3.8.1 Canonical Units	228
Chapter 4	Observations		233
<b></b>	4.1	Introduction	233
	4.1		234
	4.2	Obtaining Data	237
		4.2.1 Quantity of Data 4.2.2 Types of Data	238
		4.2.3 Example Applications	239
	4.3	Introduction to Sensor Systems	239
	4.4	Observation Transformations	244
		4.4.1 Geocentric Right Ascension and Declination	246
		4.4.2 Topocentric Right Ascension and Declination	248
		4.4.3 Azimuth-Elevation	250
		<ul><li>4.4.4 Practical Az-El Conversions</li><li>4.4.5 Transformations for Ecliptic Latitude and Longitude</li></ul>	255 257
Chapter 5	Cal	lestial Phenomena	263
Chapter 3			
	5.1	Solar Phenomena	263
		5.1.1 Application: Sun Position Vector	263 267
		5.1.2 Application: Sunrise, Sunset, and Twilight Times	207

			ix
	5.2	Lunar Phenomena	271
		5.2.1 Application: Moon Position Vector	272
		5.2.2 Application: Moon Rise and Set Times	276 280
	<i>5</i> 2	5.2.3 Phases of the Moon	281
	5.3	Celestial Applications 5.3.1 Application: Planetary Ephemerides	281
		5.3.2 Eclipses	285
		5.3.3 Application: Sight and Light	291
		5.3.4 Ground Illumination	295
		5.3.5 Miscellaneous Phenomena	297
Chapter 6	Orb	oital Maneuvering	303
	6.1	Historical Background	303
	6.2	Introduction	304
	6.3	Coplanar Maneuvers	305
		6.3.1 Hohmann and Bi-elliptic Transfers	308
		6.3.2 Comparing Hohmann and Bi-elliptic Transfers	314
		6.3.3 Transfers Using the One-Tangent Burn	317
		6.3.4 General Transfers	323 323
	6.4	Noncoplanar Transfers	323
		6.4.1 Introduction 6.4.2 Inclination-Only Changes	328
		6.4.3 Changes in the Longitude of Ascending Node	331
		6.4.4 Changes to Inclination and Longitude of Ascending Node	334
	6.5	Combined Maneuvers	336
		6.5.1 Minimum-Inclination Maneuvers	336
		6.5.2 Fixed-Dv Maneuvering	339
	6.6	Circular Rendezvous	343
		6.6.1 Circular Coplanar Phasing	343 350
		6.6.2 Circular Noncoplanar Phasing	356
	6.7	Continuous-Thrust Transfers	357
		6.7.1 Introduction 6.7.2 Orbit Raising	360
		6.7.3 Low-Thrust, Noncoplanar Transfers	366
	6.8	Relative Motion	372
	0.0	6.8.1 Position Solutions for Nearly Circular Orbits	377
		6.8.2 Trend Analysis	381
		6.8.3 Extending the Results	395
Chapter 7	Ini	tial Orbit Determination	401
-	7.1	Historical Background	401
	7.2		404
	1.2	7.2.1 Application: SITE-TRACK	404
	7.3		411
	,.5	7.3.1 Laplace's Method	413
		7.3.2 Gauss's Technique	417
		7.3.3 Double r-iteration	422
	7.4	Mixed Observations	427
		7.4.1 Range and Range-Rate Processing	428 430
		7.4.2 Range-only Processing	430

	7.5	Three Position Vectors and Time	432
		7.5.1 Gibbs Method	432
		7.5.2 Herrick-Gibbs	439 445
	7.6	1 WO I OSITION VECTOR'S AND TIME DAMPORT STREET	443 447
		7.6.1 Lambert—Minimum Energy 7.6.2 Lambert—Gauss's Solution	454
		7.6.3 Lambert—Universal Variables	459
		7.6.4 Lambert Solution—Battin Method	464
	7.7	Application: Targeting Problem	468
Chapter 8	Spe	cial Perturbation Techniques	489
	8.1	Historical Background	489
	8.2	Introduction to Perturbations	490
	8.3	Encke's Formulation	495
	8.4	Cowell's Formulation	497
	8.5	<del></del>	498
	0.5	8.5.1 Implementing an Integrator and Determining Step Size	506
	8.6	Disturbing Forces	508
	0.0	8.6.1 Gravity Field of a Central Body	509
		8.6.2 Atmospheric Drag	521
		8.6.3 Third-Body Perturbations	539 543
		8.6.4 Solar-Radiation Pressure 8.6.5 Other Perturbations	546
	8.7	Forming Numerical Solutions	548
	0.7	8.7.1 Application: Simplified Acceleration Model	550
		8.7.2 Application: Complex Acceleration Model	553
	8.8	Practical Considerations	556
		8.8.1 Validating the Propagator	556
		8.8.2 Physical Data and Sources	557
Chapter 9	Ge	neral Perturbation Techniques	565
	9.1	Historical Background	565
	9.2	Introduction	569
		9.2.1 The Method of Perturbations	573
	9.3	Variation of Parameters	575
		9.3.1 Lagrangian VOP (Conservative Effects)	577
		9.3.2 Gaussian VOP (Nonconservative and Conservative Effects)	584 593
	9.4		596
	9.5	Disturbing-Potential Formulations	590 597
		<ul><li>9.5.1 Gravity Potential in Terms of the Satellite's Orbital Elements</li><li>9.5.2 Third-Body Potential in Terms of the Satellite's Orbital Elements</li></ul>	
		<ul><li>9.5.2 Third-Body Potential in Terms of the Satellite's Orbital Elements</li><li>9.5.3 Tidal-Motion Potential in terms of the Satellite's Orbital Element</li></ul>	s 601
	9.6	1700	602
	7.0	9.6.1 Central-Body Analysis	602
		9.6.2 Drag Analysis	626
		9.6.3 Third-Body Analysis	634 640
		9.6.4 Solar-Radiation Analysis	040

	9.7	Forming Analytical Solutions 9.7.1 Application: Perturbed Two-body Propagation 9.7.2 Kozai's Method 9.7.3 Brouwer's Method 9.7.4 Operational Applications	644 645 647 649 651
	9.8	Semianalytical Solutions	652
	0.0	9.8.1 The Draper Semianalytical Satellite Theory (DSST)	653
	9.9	Practical Considerations	658
		9.9.1 Expected Accuracy of Propagation Techniques 9.9.2 Initial Conditions, Sources, and Conversions	658 663
	9.10	Summary of Perturbation Effects	667
Chapter 10	Orb	it Determination and Estimation	673
	10.1	Historical Background	673
		Linear Least Squares	676
		10.2.1 Error Analysis	681
		10.2.2 Linear Weighted Least Squares	691
	10.3	Nonlinear Least Squares	693
	10.4	Application: Orbit Determination With	<b>704</b>
		Differential Correction	701 703
		10.4.1 Calculating the Partial-Derivative Matrix (A) 10.4.2 Implementing Least-Squares Techniques	703 707
	10.5	Sequential-Batch Least Squares	713
		Kalman Filtering	718
	10.0	10.6.1 Kalman Filter (Linear System)	726
		10.6.2 Linearized Kalman Filter (LKF)	731
		10.6.3 Extended Kalman Filter (EKF)	733
		10.6.4 Successful Filtering Applications	736
	10.7	10.6.5 Summary of Kalman Filters  Forming Differential Correction Solutions	737 739
	10.7	Forming Differential Correction Solutions	739
		10.7.1 Calculating the Matrices 10.7.2 Implementing Differential-Correction Solutions	750
	10.8	Practical Considerations	753
	10.0	10.8.1 Data	754
		10.8.2 Update Intervals and Fit Spans	757
		10.8.3 Analyzing the Results of Differential Corrections	759
		10.8.4 Application: Detecting Maneuvers 10.8.5 Application: Predicting Reentry	762 762
Chapter 11	Miss	sion Analysis	765
	11.1	Introduction	765
		11.1.1 Satellite Populations	767
	11.2	Mission Orbits	768
		11.2.1 Stationkeeping	771 773
	11 2	11.2.2 Mission Planning Geometries for Surveillance and Reconnaissance	774
		Designing and Maintaining Mission Orbits	781
	11.4	11.4.1 Sun-Synchronous Orbits	782
		11.4.2 Repeat-Groundtrack Orbits	788

	11.4.3 Minimum Altitude Variation Orbits	797
	11.4.4 Frozen-Orbit Eccentricity and Argument of Perigee	803
		805
11.5	•	812
		813 815
		817
11.6		825
11.0	11.6.1 PREDICT Formulation	825
	11.6.2 Rise / Set	830
11.7	Determining Close Approaches	835
	<ul><li>11.7.1 Finding the Close-Approach Functions</li><li>11.7.2 Statistical Analysis</li></ul>	837 845
Dict	ionary of Symbols	851
Modeling the Atmosphere		865
B.1	Jacchia-Roberts Atmosphere	865
		865
		867
	B.1.3 Evaluating Density	868
	B.1.4 Robert's Corrections to Density	869
B.2	Russian GOST Atmosphere	874
Mat	thematical Fundamentals	881
C.1	Introduction	881
C.2	Vector Fundamentals	881
C.3	Matrix Fundamentals	883
C.4	Trigonometric Fundamentals	884
	C.4.1 Planar Trigonometry	886
	C.4.2 Spherical Trigonometry	887
C.5	Numerical Techniques	890
	C.5.1 Polynomial Solutions	890
	C.5.2 Interpolation	893
	C.5.3 Blending and Splining Techniques	894
Cor	nstants and Expansions	903
D.1	Gravitational Coefficients	903
		905
D.3	Reduction Coefficients	908
D.4	Planetary Ephemerides	911
D.5	Data Sources	915
D.6	Computer Programming	917
	11.6 11.7 Dict Moc B.1 B.2 Mat C.1 C.2 C.3 C.4 C.5	11.4.5 Application: Designing a Specialized Orbit 11.5 Navigation—the Global Positioning System 11.5.1 Historical Background 11.5.2 System Introduction 11.5.3 Signals from GPS Satellites 11.6 Predicting Satellite Look Angles 11.6.1 PREDICT Formulation 11.6.2 Rise / Set 11.7 Determining Close Approaches 11.7.1 Finding the Close-Approach Functions 11.7.2 Statistical Analysis  Dictionary of Symbols  Modeling the Atmosphere  B.1 Jacchia-Roberts Atmosphere B.1.1 Evaluating Temperature B.1.2 Robert's Corrections to Temperature B.1.3 Evaluating Density B.1.4 Robert's Corrections to Density B.2 Russian GOST Atmosphere  Mathematical Fundamentals C.1 Introduction C.2 Vector Fundamentals C.3 Matrix Fundamentals C.4 Trigonometric Fundamentals C.4 Trigonometric Fundamentals C.5.1 Polynomial Solutions C.5.2 Interpolation C.5.3 Blending and Splining Techniques C.5.3 Interpolation C.5.3 Reduction Coefficients D.1 Gravitational Coefficients D.2 Planetary Constants D.3 Reduction Coefficients D.4 Planetary Ephemerides D.5 Data Sources

# **Algorithm Summary**

Algorithm 1	Find $c_2 c_3$	71
Algorithm 2	KepEqtnE	73
Algorithm 3	KepEqtnP	77
Algorithm 4	KepEqtnH	79
Algorithm 5	n to Anomaly	85
Algorithm 6	Anomaly to n	85
Algorithm 7	KEPLERCOE	89
Algorithm 8	Kepler	101
Algorithm 9	ELORB	120
Algorithm 10	RANDV	125
Algorithm 11	FINDTOF	132
Algorithm 12	IJK TO LATLON	177
Algorithm 13	IJK TO LATLON	178
Algorithm 14	Julian Date	186
Algorithm 15	LSTIME	192
Algorithm 16	CONVTIME	197
Algorithm 17	DMSTORAD	199
Algorithm 18	RADTODMS	199
Algorithm 19	HMSTORAD	200
Algorithm 20	RADTOHMS	200
Algorithm 21	TIMETOHMS	201
Algorithm 22	JDToGregorianDate	204
Algorithm 23	Non-rotating Origin	211
Algorithm 24	FK5 REDUCTION	222
Algorithm 25	Geocentric Radec	248
Algorithm 26	Topocentric	249
Algorithm 27	RAZEL	254
Algorithm 28	AZELTORADEC	257
Algorithm 29	Sun	265
Algorithm 30	SUNRISESET	269
Algorithm 31	Moon	274
Algorithm 32	MOONRISESET	276
Algorithm 33	PLANETRV	283
Algorithm 34	Shadow	287
Algorithm 35	SIGHT	294
Algorithm 36	Hohmann Transfer	311
Algorithm 37	Bi-elliptic Transfer	312
Algorithm 38	One-Tangent Burn	319
Algorithm 39	Inclination Only	330