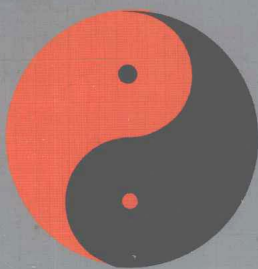


NIELS BOHR

CENTENNIAL CONFERENCES 1985



SEMICLASSICAL DESCRIPTIONS OF ATOMIC AND NUCLEAR COLLISIONS

Editors
Jens BANG
Jorrit DE BOER

SEMICLASSICAL DESCRIPTIONS OF ATOMIC AND NUCLEAR COLLISIONS

N667
1985

*Proceedings of the Niels Bohr Centennial Conference
Copenhagen, March 25–28, 1985*

Edited by

Jens BANG

Associate Professor of Physics
Niels Bohr Institute
Copenhagen, Denmark

Johrit DE BOER

Professor of Physics
University of Munich
Munich, F.R.G.

1985

NORTH-HOLLAND
AMSTERDAM • OXFORD • NEW YORK • TOKYO

© Elsevier Science Publishers B.V., 1985

All rights reserved. No part of this publication may be reproduced, stored in a retrieval system, or transmitted, in any form or by any means, electronic, mechanical, photocopying, recording or otherwise, without the prior permission of the publisher, Elsevier Science Publishers B.V. (North-Holland Physics Publishing Division), P.O. Box 103, 1000 AC Amsterdam, The Netherlands. Special regulations for readers in the USA: This publication has been registered with the Copyright Clearance Center Inc. (CCC), Salem, Massachusetts. Information can be obtained from the CCC about conditions under which photocopies of parts of this publication may be made in the USA. All other copyright questions, including photocopying outside of the USA, should be referred to the publisher.

ISBN: 0 444 86972 7

Published by:

North-Holland Physics Publishing
a division of
Elsevier Science Publishers B.V.
P.O. Box 103
1000 AC Amsterdam
The Netherlands

Sole distributors for the U.S.A. and Canada:

Elsevier Science Publishing Company, Inc.
52 Vanderbilt Avenue
New York, N.Y. 10017
U.S.A.

Library of Congress Cataloging-in-Publication Data

Main entry under title:

Semiclassical descriptions of atomic and nuclear
collisions.

Includes indexes.

1. Collisions (Nuclear physics)--Congresses.

I. Bang, Jens. II. De Boer, Jorrit, 1930-
QC794.6.C6S45 1985 539.7'54 85-15594
ISBN 0-444-86972-7 (Elsevier Science)

Printed in the Netherlands

PREFACE

The connection between quantum theory and classical physics was always one of the main themes in the work of Niels Bohr. One may think of the correspondence principle which played such a central role in his - and his pupils' - development of quantum physics, but also of his later works, where he introduced semiclassical methods in the treatment of atomic penetration and e.g. of nuclear fission. It was therefore felt appropriate to use this subject for the first in a series of conferences held celebrating the Niels Bohr Centennial.

Recent years have seen a very fruitful development in the theory of the classical limit of quantum theory. This has had an impact on the description of a large group of phenomena. The papers presented at this symposium deal with theoretical aspects of these approximations as well as applications to nuclear and atomic or molecular collisions, including chemical reactions and penetration phenomena. In these fields also experimental research, partly inspired by the semiclassical descriptions, has recently developed in interesting directions.

It was therefore natural to take the symposium as an occasion to bring together theoreticians and experimentalists from nuclear and atomic physics to promote a fruitful cross-breeding of ideas and experience in these fields.

The proceedings contain the manuscripts submitted by the invited speakers in the sequence in which they were presented at the conference. F. Zachariasen was unable to attend the symposium, but his manuscript is included. We are especially grateful to those contributors (G. Billing, P. Braun-Munzinger, M.S. Child and J.P. Dahl) who were kind enough to produce their articles at very short notice.

We wish to thank the staff members of the Niels Bohr Institute for their help in organizing this conference and Mary Carpenter (North-Holland Physics Publishing) for her help in preparing the camera-ready copy.

Jens Bang
Jorrit de Boer
Copenhagen, May 1985

ACKNOWLEDGEMENT

The organizers wish to thank the Danish National Science Research Council, the Nordic Committee for Accelerator Based Research and Nordita for financial support. Substantial funds by the Ib Henriksen Foundation and from Hafnia Hånd i Hånd are gratefully acknowledged.

J. BANG
Niels Bohr Institute
Copenhagen University
Denmark

B. MOTTELSON
NORDITA
Copenhagen
Denmark

L. KOEBACH
Institute of Physics
University of Bergen
Norway

J.M. HANSTEEN
Institute of Physics
University of Bergen
Norway

J. DE BOER
Department of Physics
University of Munich
West Germany

K. TAULBJERG
Institute of Physics
Aarhus University
Denmark

TABLE OF CONTENTS

PREFACE	v
ACKNOWLEDGEMENT	vi
On the History of the so-called WKB-Method from 1817 to 1926 N. FRÖMAN and P.O. FRÖMAN	1
Semiclassical Methods in Chemical Dynamics W.H. MILLER	9
ELASTIC AND INELASTIC PROCESSES	
Low-Energy Atomic and Molecular Collisions J.P. TOENNIES	29
Inner-Shell Coulomb Ionization E. MERZBACHER	49
Variational Methods in Inner-Shell Ionization M. KLEBER	65
Nuclear Elastic Scattering P. BRAUN-MUNZINGER	85
Absorptive Potentials G. POLLAROLO	87
Coulomb Excitation Th.W. ELZE	101
Semiclassical Functional-Integral Methods for Few and Many-Body Systems S. LEVIT	119

REARRANGEMENT COLLISIONS

Nuclear Transfer Reactions A. WINTHER	137
Semiclassical Aspects of Transfer Reactions P.D. BOND	151
Transfer Involving Deformed Nuclei J.O. RASMUSSEN, M.W. GUIDRY and L.F. CANTO	167
The Theory of Electron Capture J.S. BRIGGS	183
Electron Capture from the K-Shell: Experimental C.L. COCKE	205
Observation of the Thomas Peak E. HORSDAL	227
WKB Approximations D.M. BRINK	241
COMPLEX COLLISIONS	
Nuclear Transport-Phenomena in Low-Energy Heavy-Ion Collisions J.R. HUIZENGA and W.U. SCHRÖDER	255
Friction in Nuclear Dynamics W.J. SWIATECKI	281
Fusion Reactions at Low Energy M. BECKERMAN	315
Uniform Approximations and Condon Reflection Effects M.S. CHILD	333
A Semiclassical Approach to Molecular Energy-Transfer G.D. BILLING	339
Multiple Electron Capture R. MANN	347
WKB Approach to Nonseparable Wave Equations U. FANO	367

The Phase-Space Representation of Quantum Mechanics and the Bohr-Heisenberg Correspondence Principle J.P. DAHL	379
Quantum Phase-Space Distributions and Quantum Transport-Theory F. ZACHARIASEN	395
PENETRATION OF ATOMIC PARTICLES THROUGH MATTER	
The Barkas Effect and other Higher-Order Z_1 - Contributions to the Stopping Power H.H. ANDERSEN	409
Heavy-Ion Stopping H. GEISSEL	431
Semiclassical Description of Scattering by Atoms and by Atomic Strings J.U. ANDERSEN	463
INDEXES	
AUTHOR INDEX	483
SUBJECT INDEX	485

ON THE HISTORY OF THE SO-CALLED WKB-METHOD FROM 1817 TO 1926

Nanny FRÖMAN and Per Olof FRÖMAN

Institute of Theoretical Physics, University of Uppsala, Thunbergsvägen 3,
S-752 38 Uppsala, Sweden

The mathematical approximation method which, since the breakthrough of quantum mechanics, is usually called the WKB-method, has really been known for a very long time. The method describes various kinds of wave motion in an inhomogeneous medium, where the properties change only slightly over one wavelength, and also provides the connection between classical mechanics and quantum mechanics. To a surprisingly large extent it can be found already in an investigation by Carlini (1817) on the motion of a planet in an unperturbed elliptic orbit. After that the method was independently developed and used by many people. The important connection formulas were, however, missing until Rayleigh (1912) very implicitly and Gans (1915) somewhat more explicitly derived one of them, later rediscovered independently by Jeffreys (1923), who also derived another connection formula (although not in quite correct form), and by Kramers (1926).

In 1817 Carlini¹ treated an important problem in celestial mechanics. He considered the motion of a planet in an elliptic orbit around the sun, with the perturbations from all other gravitating bodies neglected. Using a polar co-ordinate system in the plane of the planetary motion, with the origin at the sun, one can then express the polar angle as $2\pi t/T$ plus an infinite series containing sines of multiples of $2\pi t/T$, where t is the time counted from a perihelion position, i.e. from a moment when the planet is closest to the sun, and T is the time for one revolution of the planet in its orbit. The problem treated by Carlini was to determine the asymptotic behaviour of the coefficients of the sines in this series for large values of the summation index. In his treatment of this problem Carlini had to investigate a function s of a variable x . This function, which Carlini defined by a power series in x , is proportional to the function which is now called a Bessel function of the first kind, with the index p and the argument proportional to px . Carlini, who needed a useful approximate formula for this function when its argument is smaller than its order p , which tends to infinity, showed that $s(x)$ satisfies a linear, second-order differential equation containing the large parameter p . In this differential equation Carlini introduced a new dependent variable y by putting

$$s = \exp\left(\frac{1}{2}p \int^x y \, dx\right) . \quad (1)$$

Then he expanded the function y in inverse powers of p . When he introduced this expansion into the differential equation for y and identified terms containing the same power of $1/p$, Carlini obtained recursive formulas which give what is now usually called the WKB-approximation, with higher-order terms included, for the solution of the differential equation satisfied by $s(x)$. In explicit form he gave essentially the second-order WKB-approximation for the solution in a classically forbidden region. If we express Carlini's result for the function $s(x)$ in terms of the Bessel function $J_p(\xi)$, where ξ is proportional to p , we obtain the result

$$J_p(\xi) = \frac{\xi^p}{2^p p!} \exp\left\{p \left[\left(1 - \frac{\xi^2}{p^2}\right)^{\frac{1}{2}} - 1 - \ln \frac{1 + (1 - \xi^2/p^2)^{\frac{1}{2}}}{2} \right] - \frac{1}{2} \ln \left(1 - \frac{\xi^2}{p^2}\right)^{\frac{1}{2}} + \frac{1}{p} \left[\frac{1}{12} + \frac{1}{8 (1 - \xi^2/p^2)^{\frac{1}{2}}} - \frac{5}{24 (1 - \xi^2/p^2)^{\frac{3}{2}}} \right] + \dots \right\} \quad (2)$$

where $0 \leq \xi < p$, ξ/p is independent of p , and $p (>0)$ is large. Formula (2) is essentially equivalent to the next lowest order of the asymptotic formula derived in 1909, i.e. almost a century later, by Debye¹⁷ for the Bessel function $J_p(\xi)$ when $0 \leq \xi < p$, ξ/p is independent of p , and $p \rightarrow \infty$. We also remark that if one follows in essential respects Carlini's procedure to derive (2), but uses modern developments of phase-integral technique, one can in a simple way obtain asymptotic formulas, essentially equivalent to those derived by Debye with the aid of the more complicated method of steepest descents.

Using the language of quantum mechanics, one can say that in the part of the work by Carlini, which has been described above, Carlini obtained an approximate expression for the solution of the radial Schrödinger equation in the classically forbidden region, in the absence of a physical potential. Because of the way in which the large parameter p appears in the differential equation for the function $s(x)$, Carlini's solution remains valid at $x = 0$ in any order of approximation. Carlini thus achieved automatically in any order of approximation the result which Kramers²⁶ achieved in the first-order WKB-approximation by replacing $l(l+1)$ by $(l+\frac{1}{2})^2$, where l is the orbital angular momentum quantum number.

In connection with a heat conduction problem, Liouville² in 1837 treated an ordinary, linear, second-order differential equation which he transformed into a differential equation of the Schrödinger type by means of what one now calls a Liouville transformation. Then he arrived at what one in quantal language now usually calls the first-order WKB-approximation in a classically allowed region.

In 1837 Green³ considered the motion of waves in a non-elastic fluid confined in a canal with infinite extension in the x -direction and with small breadth and depth, both of which may vary slowly in an unspecified way. The problem is described by a partial differential equation which is of the second order both with respect to the coordinate x and the time t . Green obtained an approximate solution which, for the particular case when its time-dependence is described by a sine or cosine function, reduces essentially to the first-order WKB-approximation in a classically allowed region.

The famous astronomer Encke, after whom a comet is named, drew Jacobi's attention to Carlini's work, and in 1849 Jacobi⁴ published a paper concerning improvements and corrections to Carlini's work. In this paper Jacobi characterized Carlini's work as excellent and instructive, and he considered the problem treated in the main part of Carlini's publication as one of the most difficult problems of its class. Although Jacobi pointed out and corrected mistakes made by Carlini, he also pointed out that all the essential difficulties in the solution of the problem already had been vanquished by Carlini in 1817 and that Carlini's final result would have been correct if he had not made trivial mistakes in his calculations.

In 1850 Jacobi published a translation into German, with critical comments and extensions, of Carlini's investigation⁵. In this publication Jacobi again emphasized that, although the work by Carlini of 1817 contains many mistakes, and the final results are incorrect, this work, because of the method used there and the boldness of its composition, yet belongs indisputably to the most important works concerning the determination of the values of functions of large numbers. More than three decades after the original publication of Carlini's work, Jacobi thus considered it highly desirable to republish it with the necessary improvements and extensions included.

The problem in celestial mechanics, which Carlini had treated by starting from a formula given by Lagrange, was later solved more generally and in much simpler ways in 1856 by Scheibner^{6,7}, who attacked the problem from quite different starting points. In his first paper, Scheibner⁶ used a peculiar and very general method, which recommends itself by the brevity and ease of the calculations. In his second paper, Scheibner⁷ used Cauchy's powerful theory of complex integration. As an indication of the importance of Scheibner's papers we mention that, almost a quarter of a century after they had first been published, the first paper⁶, originally written in English, was republished in German translation⁸, and the second paper⁷, originally written in German, was republished in abbreviated form⁹. Scheibner thus solved the actual problem in celestial mechanics much simpler and more satisfactorily than Carlini, but

the more complicated investigation by Carlini yielded the very fundamental result which is now called the WKB-approximation of arbitrary order. We mention Scheibner's work only to illuminate Carlini's treatment. The methods used by Scheibner are otherwise not related to the history of the so-called WKB-approximation.

In the second edition, published in 1895, of his well-known book on hydrodynamics, Lamb¹⁰ treated (on pp. 291-296) the propagation of waves in a canal of gradually varying section on the basis of the investigation by Green³.

Apparently unaware of the results obtained earlier by other authors, de Sparre¹¹ derived in 1898 essentially what is now called the second-order WKB-approximation for a second-order differential equation.

From a purely mathematical point of view, Horn^{12,13} in 1899 considered, for real values of the independent variable, the asymptotic solution of a linear, second-order differential equation containing a large parameter. In 1906 and 1907 Schlesinger^{14,15} generalized Horn's mathematical investigations by treating, for complex values of the independent variable, a linear system of first-order differential equations containing a large parameter. Referring to the method used by Green³, Birkhoff¹⁶ in 1908 continued Horn's and Schlesinger's work by investigating mathematically the asymptotic character of the solutions of certain arbitrary-order linear differential equations containing a large parameter.

In a paper concerning the propagation of waves through a stratified medium, Rayleigh¹⁸ in 1912 treated the one-dimensional time-independent wave equation by writing the solution as an amplitude times a phase factor. He found the exact relation between amplitude and phase [his eq. (73)], but he did not point out the great importance of this relation, which since 1930 has been used with great success by several authors. Rayleigh obtained what is now generally known as the first-order WKB-approximation in a classically allowed region. By pursuing the approximations he also obtained the next correction to the amplitude (in approximate form) and to the phase. Considering then the case of total reflection due to a turning point, Rayleigh introduced into the wave equation a linear approximation in a certain region around the turning point and was thus able to obtain an approximate solution expressed by an Airy function in that region. When he used asymptotic approximations for the Airy function, he obtained a result which is closely related to a connection formula for the WKB-approximation.

On the basis of Maxwell's electromagnetic theory, Gans¹⁹ in 1915 treated the propagation of light in an inhomogeneous medium, where the index of refraction depends only on one cartesian coordinate. He obtained the first-order

WKB-approximation for the solution of the one-dimensional wave equation. Considering total reflection, Gans approximated the square of the index of refraction by a linear function in the region around the turning point which gives rise to the total reflection. He was thus able to express the solution on each side of the turning point approximately in terms of Hankel functions of the order $1/3$. Matching these approximate solutions at the turning point, and using asymptotic approximations for the Hankel functions on both sides of the turning point, Gans obtained a result [see in Ref. 19 eqs. (69) and p. 726] which, although not in quite explicit form, is equivalent to the connection formula for the first-order WKB-approximation which starts from the exponentially small wave function in the region into which the light penetrates only as an evanescent wave.

In a paper dealing with certain hypotheses as to the internal structure of the earth and moon, Jeffreys²⁰ in 1915 obtained (on pp. 211-213) essentially the first-order WKB-approximation for the solution of a linear, second-order differential equation.

In an investigation concerning the aerodynamics of a spinning shell, Fowler, Gallop, Lock and Richmond²¹ in 1921 treated a system of two coupled ordinary, linear differential equations containing a large parameter, one of the equations being inhomogeneous and of the second order, the other being homogeneous and of the first order. Referring to the papers by de Sparre¹¹, Horn^{12,13}, Schlesinger^{14,15} and Birkhoff¹⁶, Fowler et al.²¹ investigate the asymptotic expansion of the solution of the above-mentioned system of differential equations for large values of the parameter. In connection with this problem, the authors consider in particular a homogeneous, linear differential equation of the second order which they solve by writing the solution as the product of an amplitude and a phase factor. Finding the exact relation between amplitude and phase, they express the phase in terms of the amplitude which they obtain as an asymptotic expansion in inverse powers of the square of the large parameter. The authors make the important remark that by separating the solution correctly into the product of an amplitude and a phase factor they gain the advantage over other methods that they obtain in one step a solution with the error inversely proportional to the square of the large parameter, whereas this requires two steps in the usual procedures.

Referring to the above-mentioned previous work by Green, Lamb, Horn, Jeffreys and Fowler et al., Jeffreys²² in 1923 derived what is now usually called the WKB-approximation for the solution of an ordinary, homogeneous, linear differential equation of the second order. In the region of a turning point Jeffreys, like Rayleigh and Gans, introduced a linear approximation in

the differential equation and was thus able to express the solution there approximately in terms of Bessel functions of the order $1/3$. Using asymptotic approximations for these functions, Jeffreys obtained the previously mentioned connection formula and another connection formula which was, however, not given in quite correct form. The question of the one-directional nature of the connection formulas was not clarified until later.

Obviously unaware of the existence of the work by the previous authors, Brillouin^{23,24}, Wentzel²⁵ and Kramers²⁶ in 1926 introduced analogous considerations in quantum mechanics. Brillouin established, for a system of particles, the connection between the Schrödinger equation of quantum mechanics and the Hamilton-Jacobi equation of classical mechanics, while Wentzel and Kramers introduced for the radial Schrödinger equation the main results obtained in the course of the developments described above. These results turned out to be extremely useful in applications of the new quantum theory and became known under the name of the WKB-method. However, Brillouin, Wentzel and Kramers contributed hardly anything new to the mathematical approximation method that had already been found by previous authors, as described in this short historical review.

REFERENCES

- 1) F. Carlini, Ricerche sulla convergenza della serie che serve alla soluzione del problema di Keplero, in: Appendice all' Effemeridi Astronomiche di Milano per l'anno 1818 (Dall' Imp. Regia Stamperia, Milano, 1817) pp. 3-48.
- 2) J. Liouville, J. Math. Pures Appl. 2 (1837) 16-35.
- 3) G. Green, Trans. Cambr. Phil. Soc. 6 (1837) 457-462.
- 4) C.G.J. Jacobi, Astron. Nachr. 28 (1849) 257-270; also published in C.G.J. Jacobi's Gesammelte Werke, Band 7, herausgegeben von K. Weierstrass (Druck und Verlag von Georg Reimer, Berlin, 1891) pp. 175-188.
- 5) F. Carlini, Astron. Nachr. 30 (1850) 197-254. Translation into German and revision by Jacobi of the paper by Carlini of 1817; also published in C.G.J. Jacobi's Gesammelte Werke, Band 7, herausgegeben von K. Weierstrass (Druck und Verlag von Georg Reimer, Berlin, 1891) pp. 189-245.
- 6) W. Scheibner, Astron. J. 4 (1856) 177-182.
- 7) W. Scheibner, Berichten der Kgl. Sächs. Ges. d. Wiss. zu Leipzig, Math.-Phys. Classe 8 (1856) 40-64.
- 8) W. Scheibner, Math. Ann. 17 (1880) 531-544.
- 9) W. Scheibner, Math. Ann. 17 (1880) 545-560.
- 10) H. Lamb, Hydrodynamics, second edition (University Press, Cambridge, 1895).

- 11) M. de Sparre, Atti della R. Acc. dei Lincei (Ser. V). Rendiconti. Classe di sc. fis., math. e nat. 7:2 (1898) 111-117.
- 12) J. Horn, Math. Ann. 52 (1899) 271-292.
- 13) J. Horn, Math. Ann. 52 (1899) 340-362.
- 14) L. Schlesinger, C.R. Acad. Sci. Paris 142 (1906) 1031-1033.
- 15) L. Schlesinger, Math. Ann. 63 (1907) 277-300.
- 16) G.D. Birkhoff, Trans. Amer. Math. Soc. 9 (1908) 219-231.
- 17) P. Debye, Math. Ann. 67 (1909) 535-558.
- 18) Lord Rayleigh (J.W. Strutt), Proc. Roy. Soc. London A 86 (1912) 207-226.
- 19) R. Gans, Ann. Physik (Vierte Folge) 47 (1915) 709-736.
- 20) H. Jeffreys, Mem. Roy. Astr. Soc. 60 (Part IV) (1915) 187-217.
- 21) R.H. Fowler, E.G. Gallop, C.N.H. Lock and W.H. Richmond, Phil. Trans. Roy. Soc. London A 221 (1921) 295-387.
- 22) H. Jeffreys, Proc. Lond. Math. Soc. (Second Series) 23 (1923) 428-436.
- 23) L. Brillouin, C.R. Acad. Sci. Paris 183 (1926) 24-26.
- 24) L. Brillouin, J. Physique Radium (Série VI) 8 (1926) 353-368.
- 25) G. Wentzel, Z. f. Physik 38 (1926) 518-529.
- 26) H.A. Kramers, Z. f. Physik 39 (1926) 828-840.

SEMICLASSICAL METHODS IN CHEMICAL DYNAMICS

William H. MILLER

Department of Chemistry, University of California, and Materials and
Molecular Research Division, Lawrence Berkeley Laboratory, Berkeley,
California 94720, USA

A general semiclassical (multidimensional WKB-type) approximation to quantum mechanics is reviewed. The principal feature of the approach is that it is able to incorporate the exact classical mechanics of the system and also the quantum principle of superposition. Applications to inelastic and reactive scattering, and to statistical mechanics and reaction rates are discussed.

1. Introduction

The use of semiclassical methods in chemical and molecular problems has become so pervasive over the last twenty years that it is somewhat intimidating to attempt a short review with the present title.¹ Since a comprehensive review is impossible, I have chosen several examples which illustrate the variety of quantum phenomena that semiclassical theory is able to describe correctly.

Perhaps the most dramatic quantum phenomenon is that of quantization itself, the fact that only discrete values of the energy are allowed for a system whose motion is bounded. The grandfather of all semiclassical descriptions of quantization is the Bohr model of the hydrogen atom and its generalization which became known as the Old Quantum Theory.² For a diatomic molecule with potential function $V(r)$, for example, the Bohr-Sommerfeld (or WKB) quantization condition is

$$(n + 1/2)\pi = \int_{r_1}^{r_2} dr \sqrt{\frac{2\mu}{\hbar^2} [E - V(r) - \frac{\hbar^2 \ell(\ell+1)}{2\mu r^2}]}, \quad (1.1)$$

where (r_1, r_2) are the two classical turning points in the effective potential ($V(r)$ plus the centrifugal term). Eq. (1.1) is an implicit equation for the energy levels $E(n, \ell)$ in terms of the vibrational and rotational quantum numbers.

For molecular systems, i.e., diatomic molecules, Eq. (1.1) is actually quite accurate: ground-state energies are given to better than 1% accuracy even for